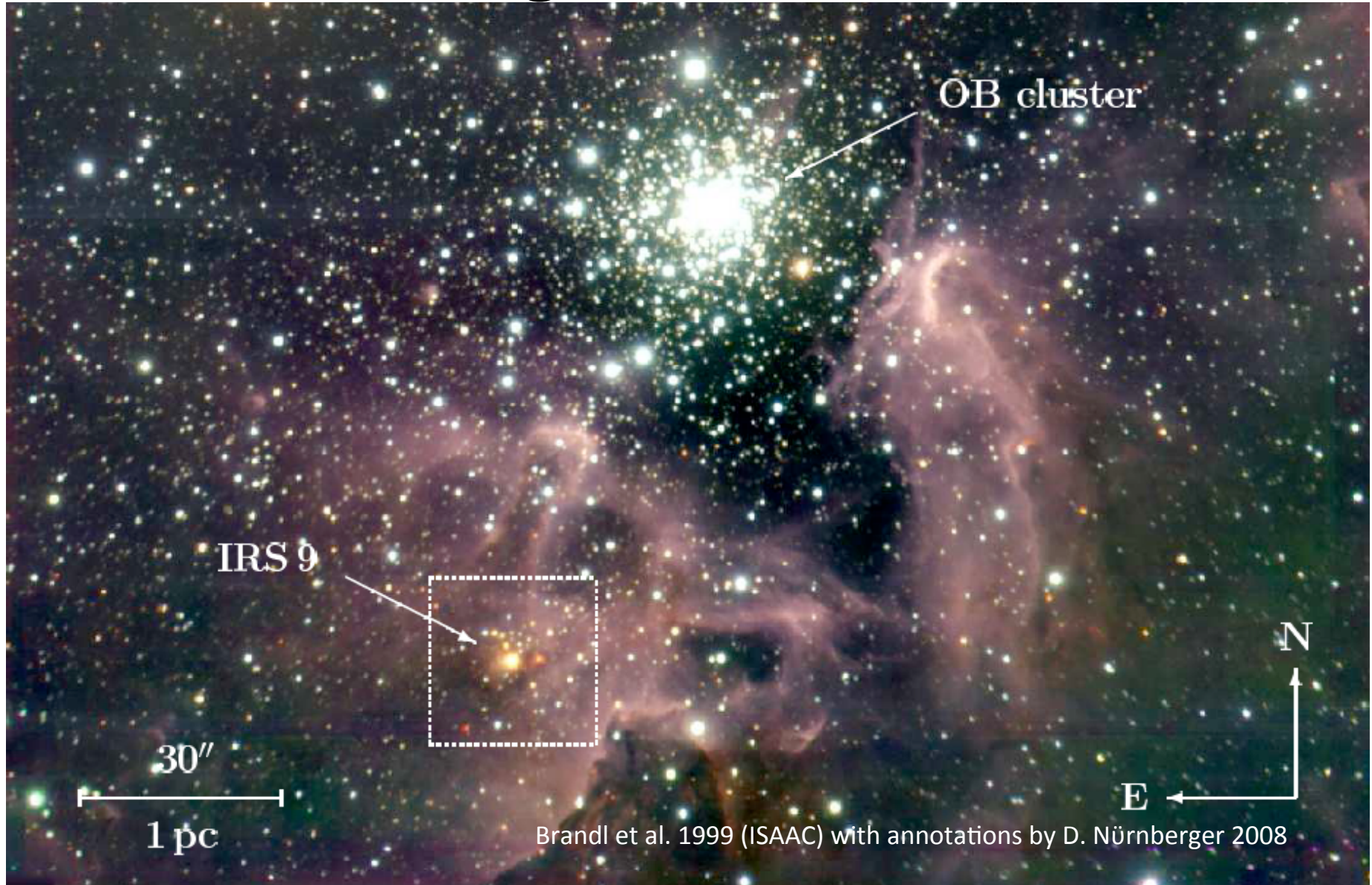


Sub-mm observations of compact cores associated with a massive young stellar object in NGC 3603

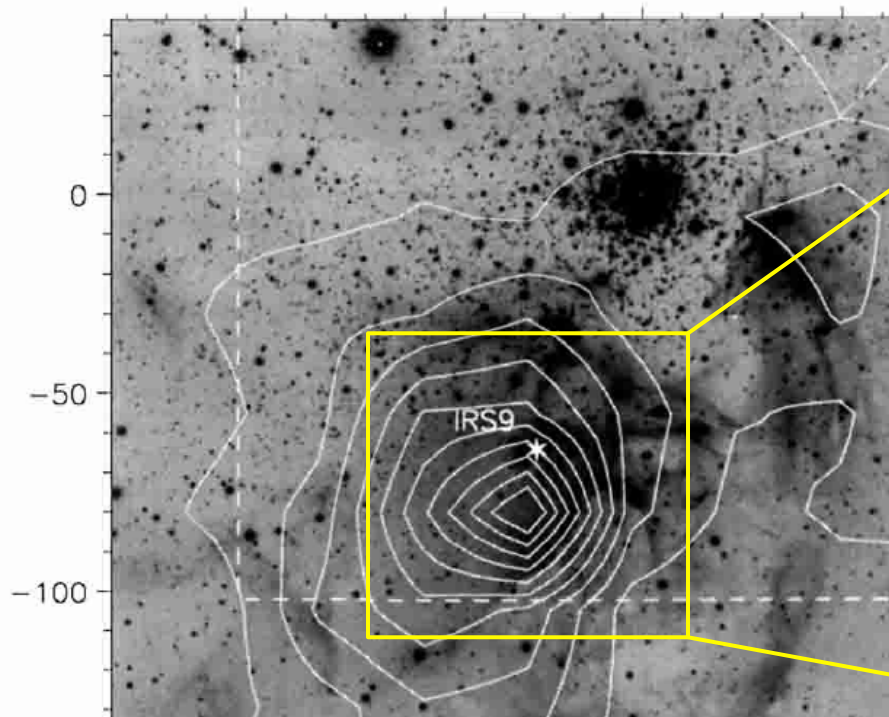
Christian Hummel (ESO), Thomas Stanke (ESO),
Roberto Galvan-Madrid (UNAM), Bärbel Koribalski (CSIRO)

HII region NGC 3603



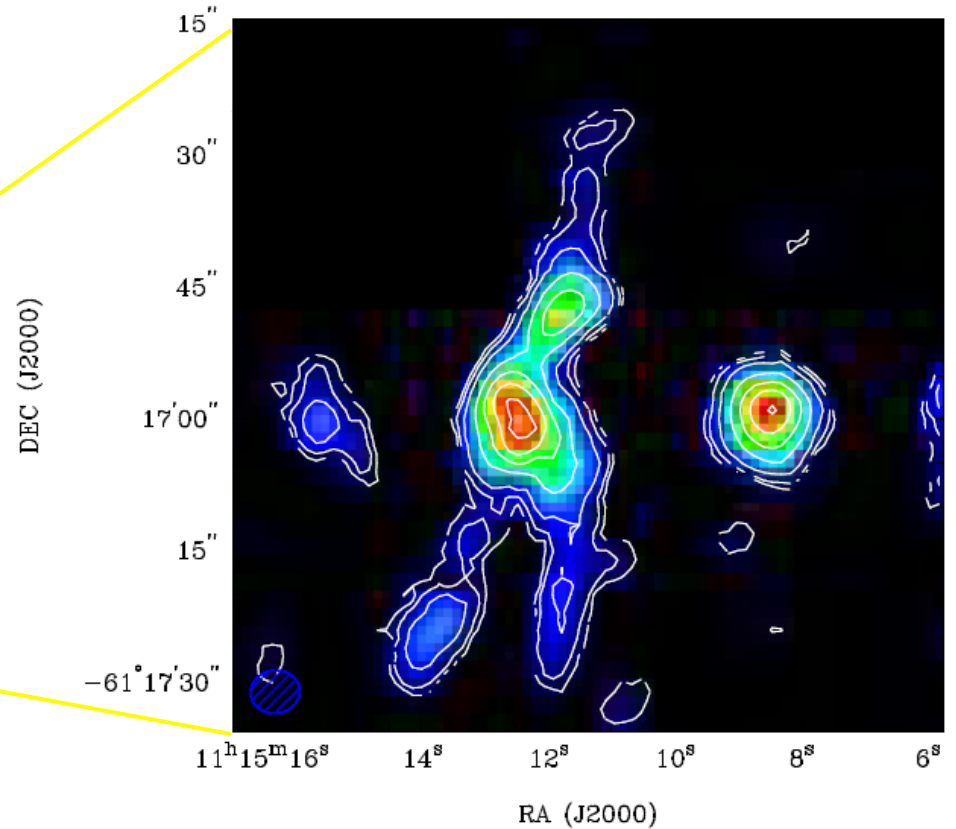
CS (2-1) high density tracer

SEST 3 mm



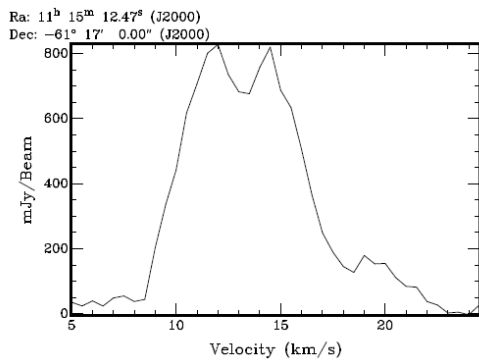
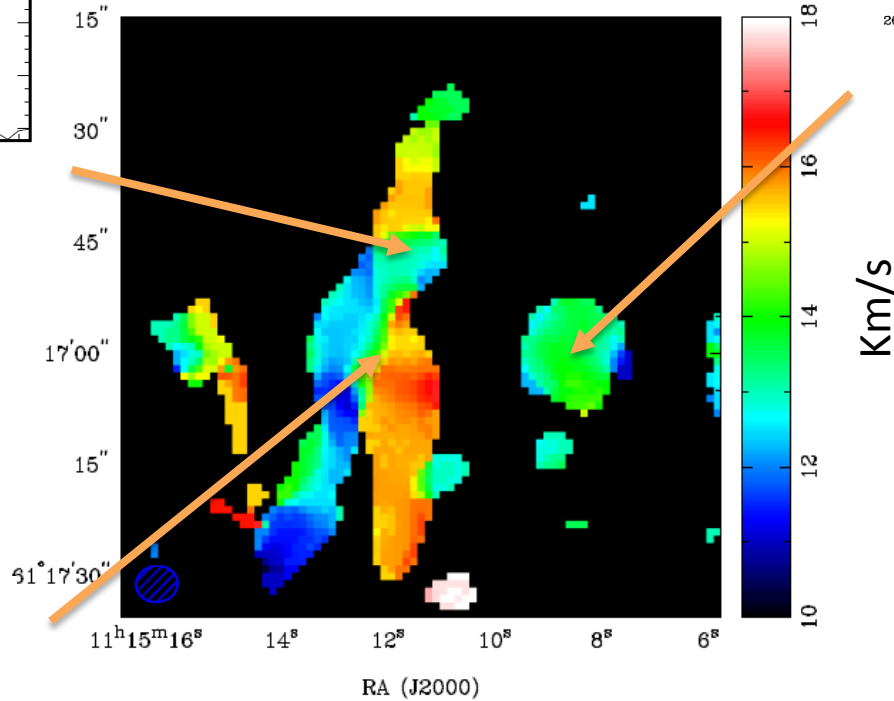
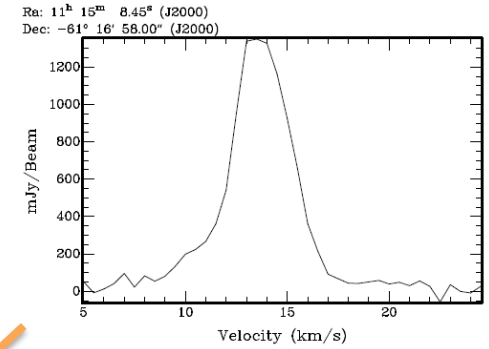
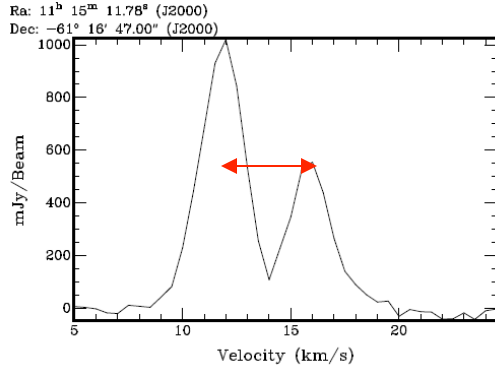
Nürnbergger and Petr-Gotzens 2002

ATCA 3 mm

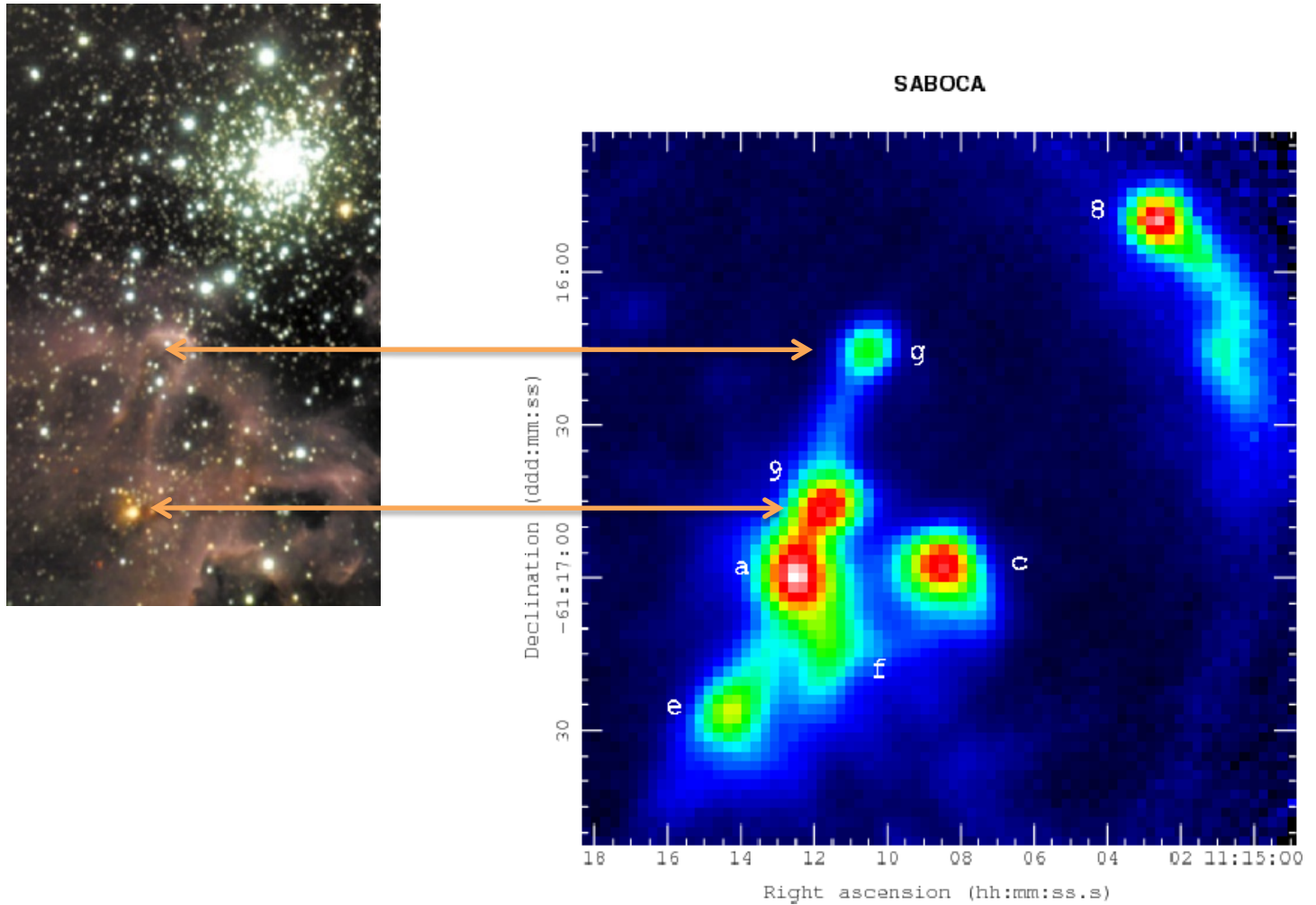


Hummel et al. (submitted)
Map by Bärbel Koribalski

Filaments



350 μm continuum w/SABOCA



Hummel et al. (submitted)

Compact core properties

	RA	Dec	F Jy	θ "	Δv km/s	M_F M_\odot	M_V M_\odot
a	11:15:12.07	-61:16:58.8	90	13.1	2.5	385	175
8	11:15:02.88	-61:15:51.5	55	9.8	–	237	–
c	11:15:08.41	-61:16:57.2	67	11.7	3.2	285	255
9	11:15:11.34	-61:16:45.2	59	10.8	2.5	254	145
e	11:15:13.73	-61:17:24.3	46	12.3	2.5	199	165
f	11:15:11.25	-61:17:12.0	45	13.6	–	192	–
g	11:15:10.33	-61:16:16.4	22	8.3	–	94	–

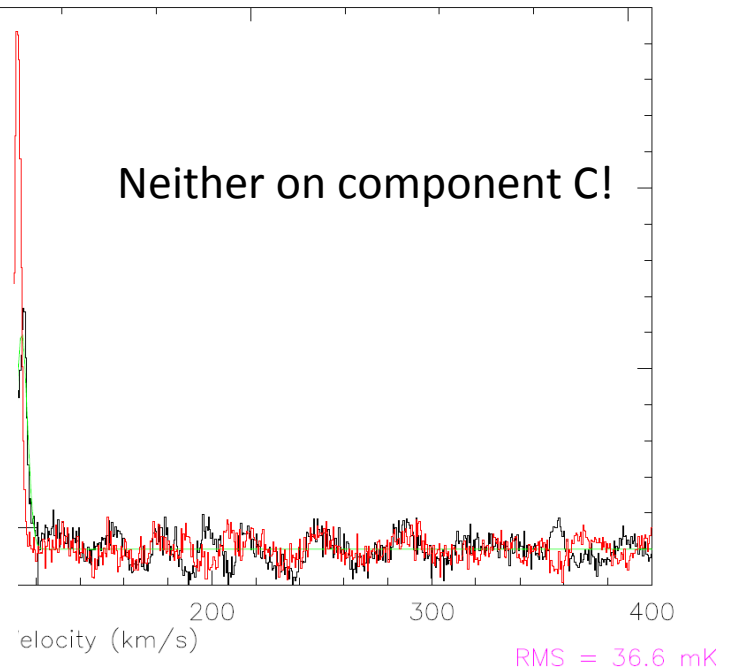
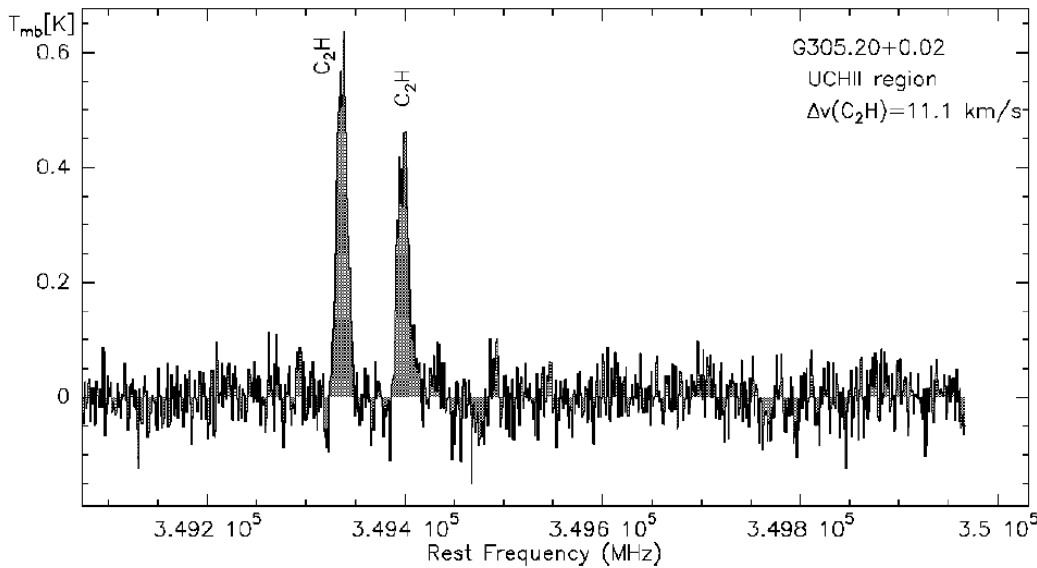
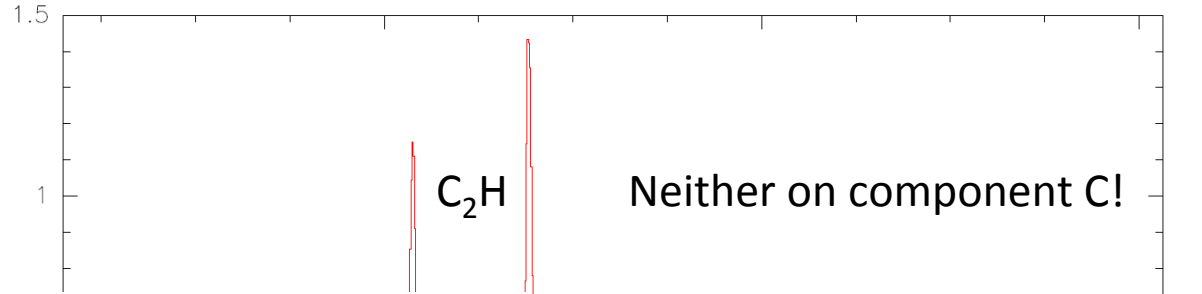
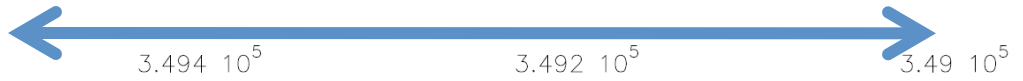
$$10'' = 0.33 \text{ pc}$$

$$\mathcal{M}_{\text{vir}} = k \cdot \mathcal{R} \cdot \Delta v_{FWHM}^2$$

SHFI sub-mm spectroscopy: aims

- Hot cores: CH_3CN (K-series of methylcyanide)
- Temperature: H_2CO (formaldehyde)
- Outflows: SiO (shock tracer), CO (3-2)
- Infall: HCO^+ and H_{13}CO^+
- HII region: $\text{H}26\alpha$, $\text{H}30\alpha$ (recombination lines)

SHeFI APEX 2

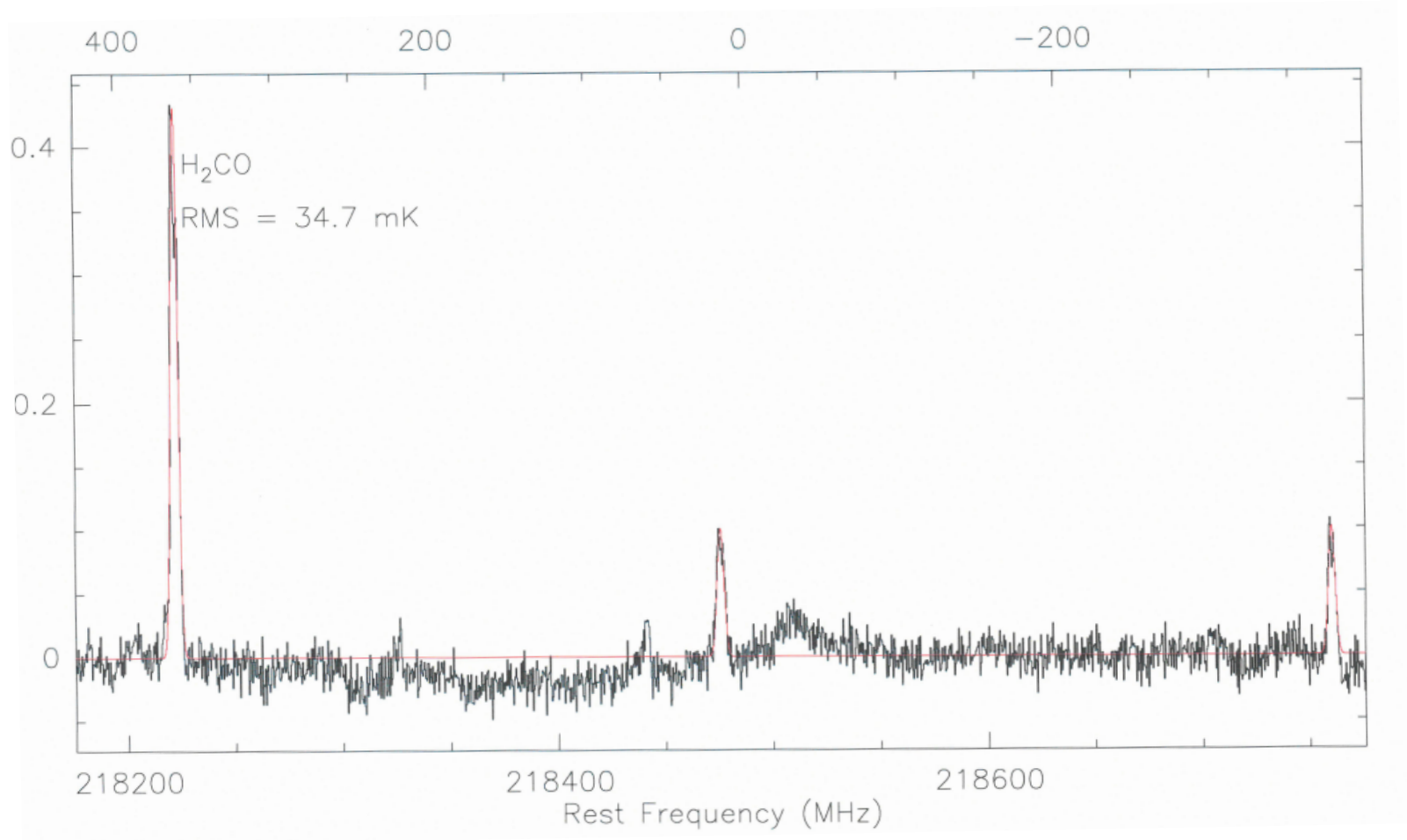


Beuther et al. 2008 (G305.20+0.02)

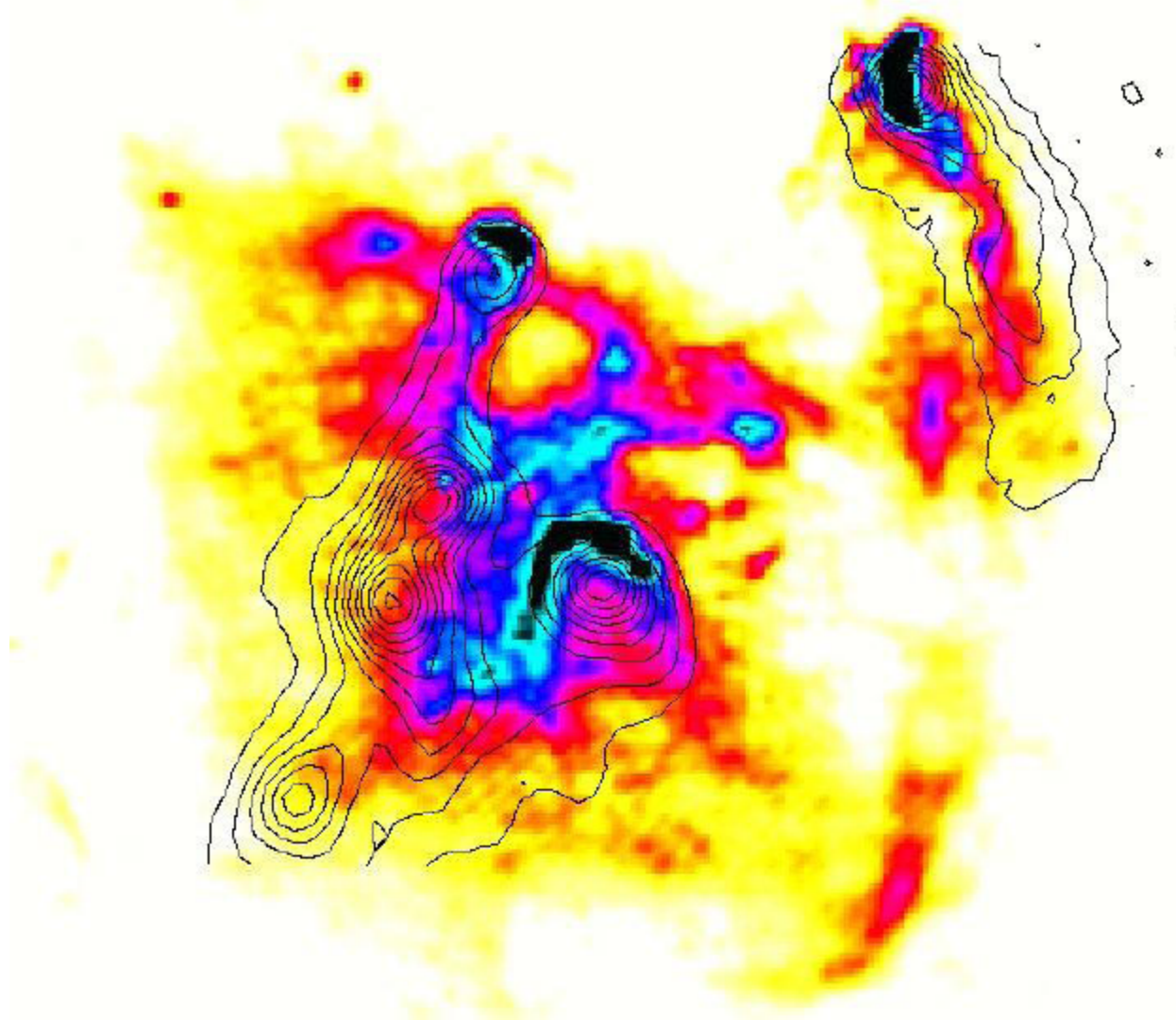
Temperature from Formaldehyde lines

45 K
 $\Delta V = 5 \text{ km/s}$

$V_{\text{LSR}} = +14 \text{ km/s}$



SABOCA 350 μm + ATCA 6 cm



...back to IRS 9A: MIR and sub-mm

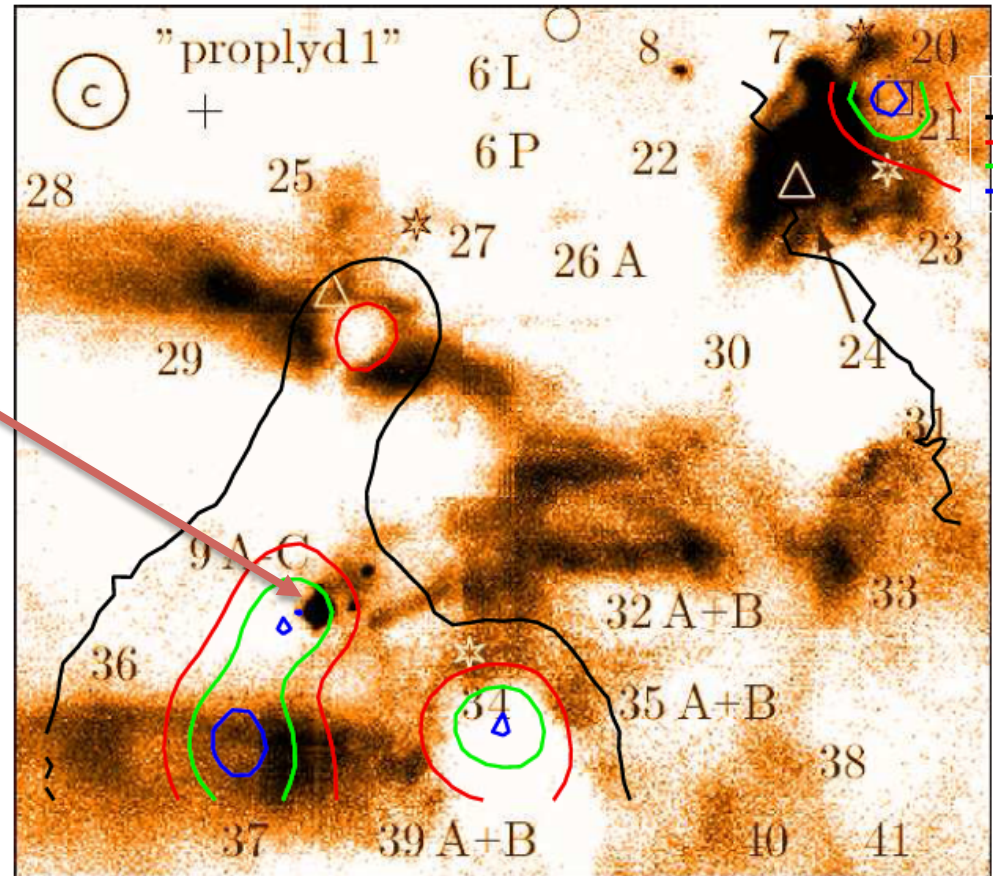
IRS 9A

D = 7 kpc

L = $10^5 L_{\odot}$

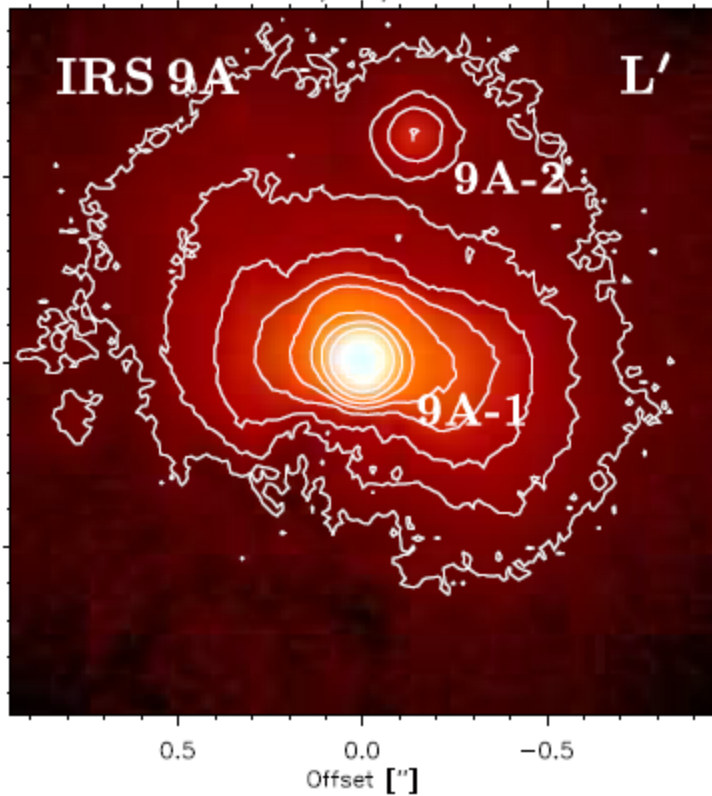
Image: 10 μm (TIMMI2)

Contours: 350 μm (SABOCA)

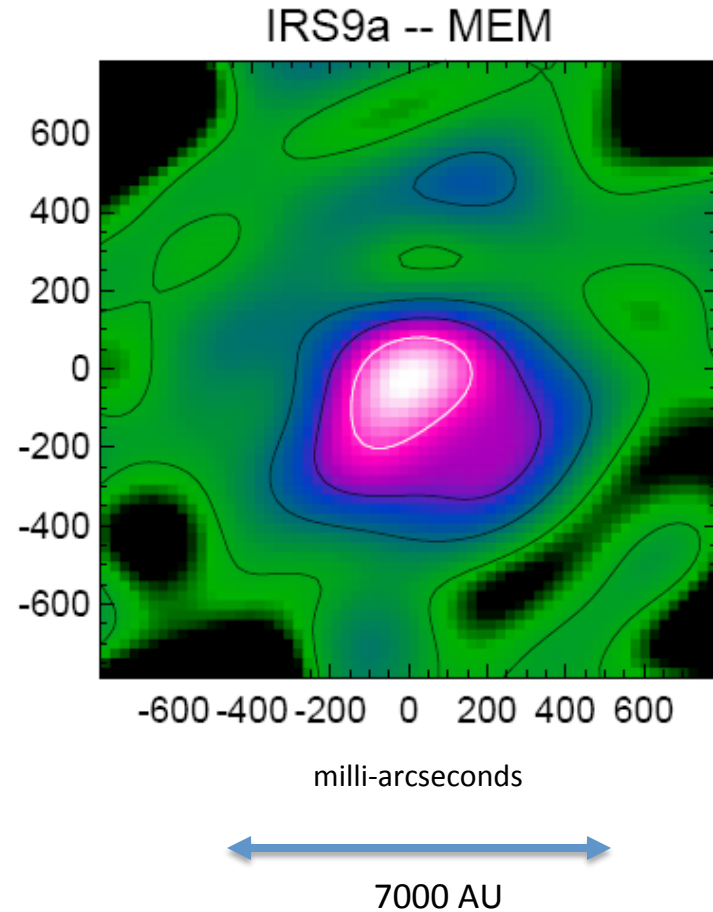


Contours: SABOCA, grey scale: Nürnberger and Stanke 2003 (TIMMI2)

The structure of the MYSO IRS 9A



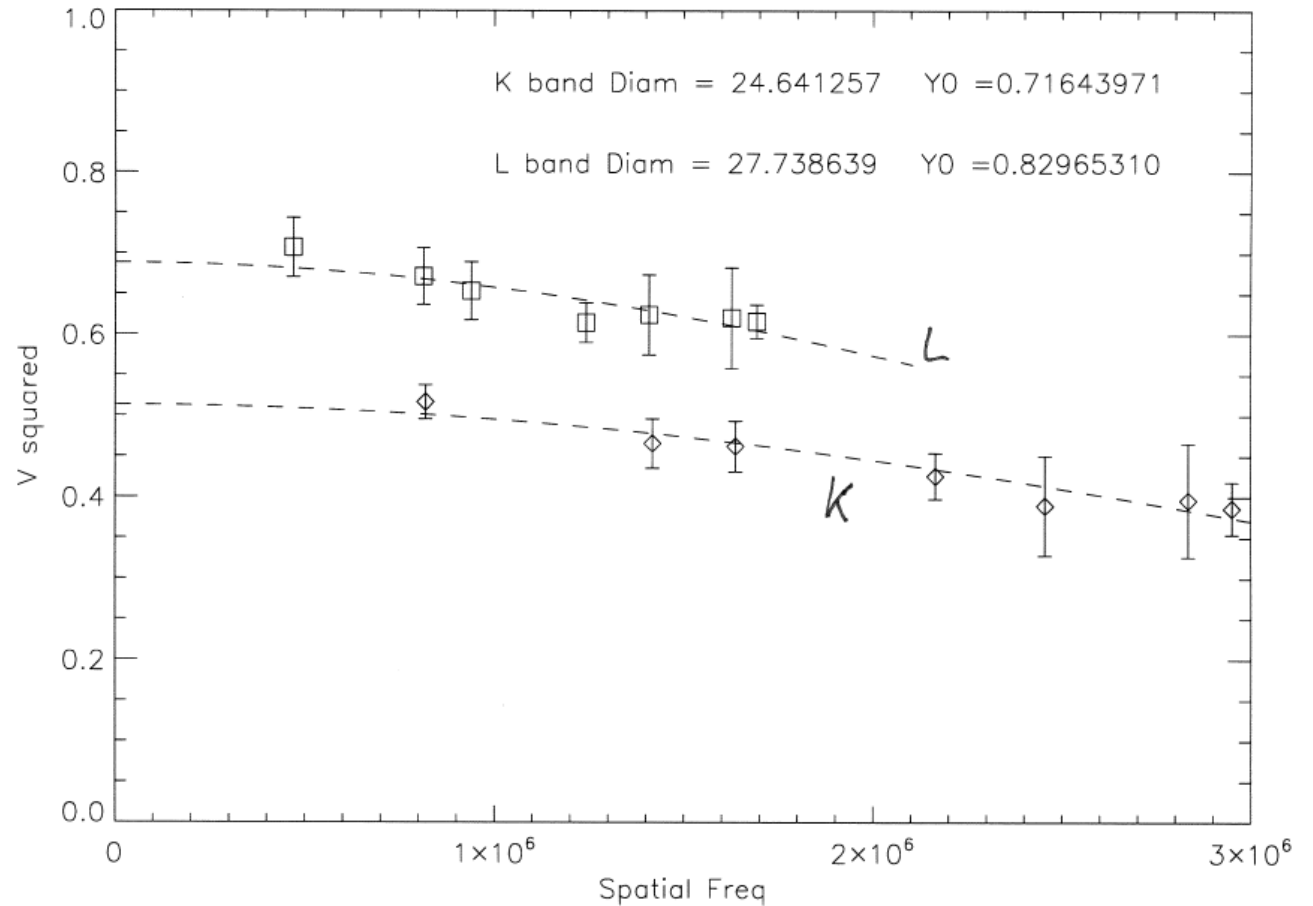
Nürnbergger 2008 (NACO @ VLT)



Vehoff et al. 2010 (T-ReCS @ Gemini S)

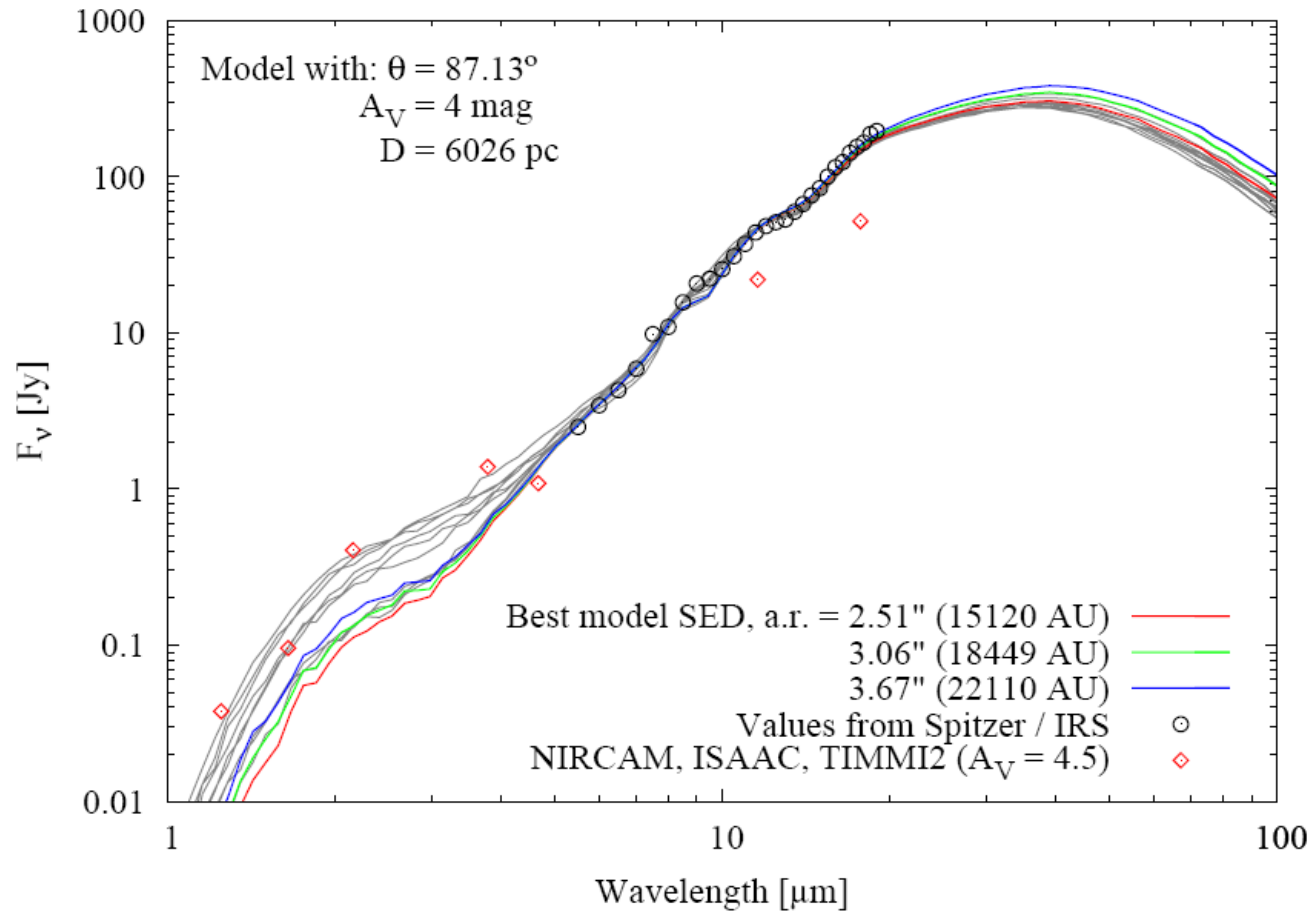
NACO Ks & Lp aperture masking

Diameters @ 7 kpc:
K band: 170 AU
L band: 190 AU



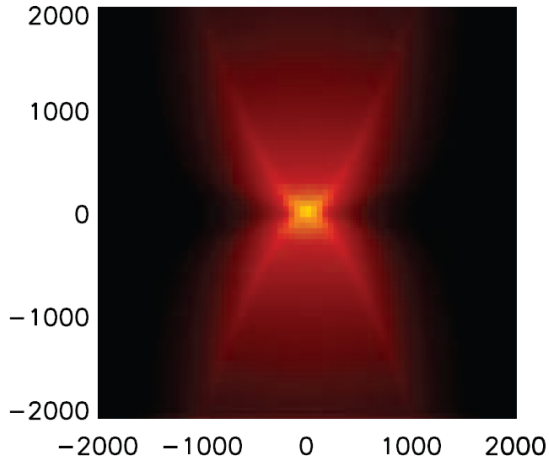
Sanchez et al. 2015 (submitted)

SEDs for disk-envelope model grid

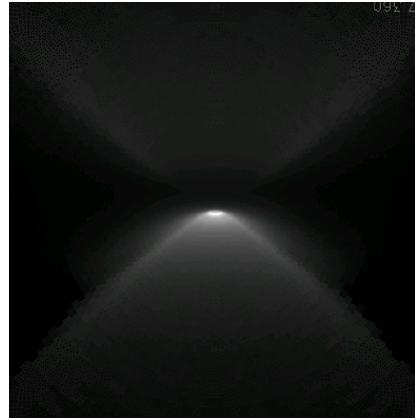


Grid by: Robitaille et al. 2006

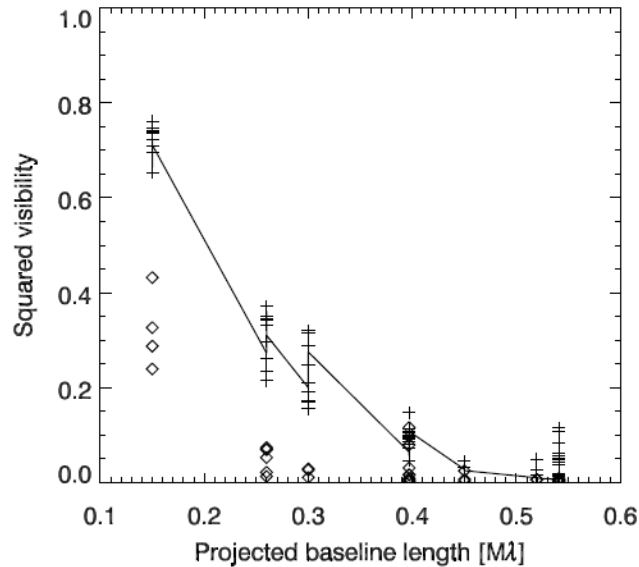
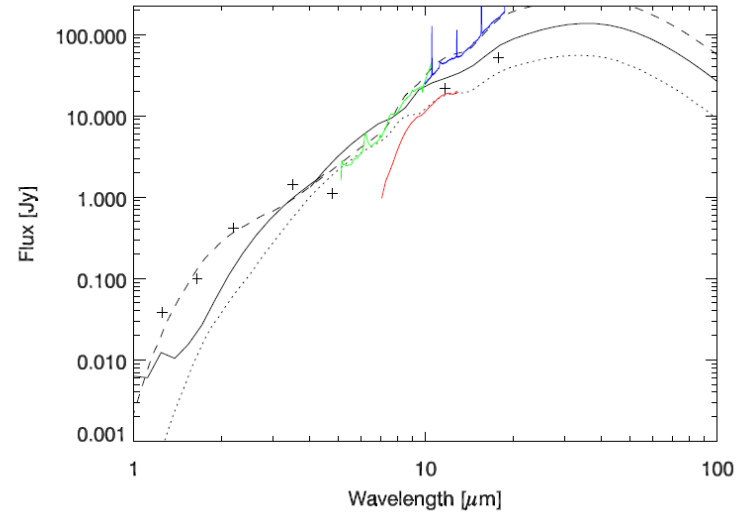
Disk-envelope models



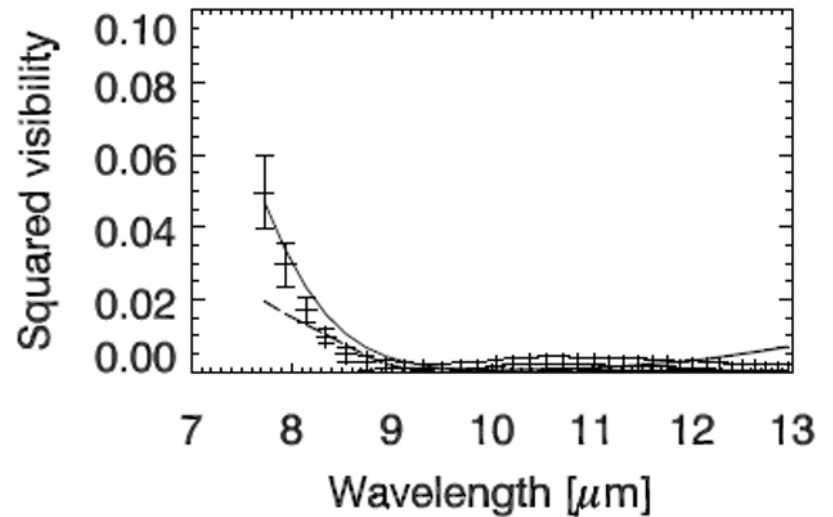
Whitney et al. 2003



Wolf 2003



Diamonds: model values



Dashed: Whitney; solid: MC3D

Summary (1)

- Massive compact cores
 - No hot cores (no methyl-cyanide lines detected)
 - Spectral features consistent with IRDCs
 - Formaldehyde lines give $T = 45$ K
 - Cores do not contain HII regions

Summary (2)

- IRS 9A
 - younger than compact cores
 - located on cluster-facing side
 - Consistent with sequential star formation scenario in NGC 3603
 - Line of sight into compact region: MIR imaging with the MATISSE interferometer at VLTI in 2017?

Detected lines

Species	Transition	ν_0 (GHz)	E_u (K)	T_c/T_9
SiO	5-4	217.104984	31.2	–
DCN	3-2	217.238531	20.9	1.32
c-HCCCH	6(1,6)-5(0,5)	217.822141	38.6	1.47
CH ₃ OH	4(2,2)-3(1,2)	218.440047	45.5	–
H ₂ CO	3(0,3)-2(0,2)	218.222188	21.0	1.39
...	3(2,2)-2(2,1)	218.475641	68.1	1.30
...	3(2,1)-2(2,0)	218.760078	68.1	1.41
C ¹⁸ O	2-1	219.5603	15.8	1.19
SO	5,6-4,5	219.949438	35.0	1.45
¹³ CO	2-1	220.3986	15.86	0.97
CO	2-1	230.538	16.6	0.75
¹³ CS	5-4	231.220688	33.3	–
H α	H30 α	231.900930	Rec.	–
SO	8,9-7,8	346.528594	78.8	2.40
H ¹³ CO+	4-3	346.998347	41.6	3.37
H ₂ CS	10(1,9)-9(1,8)	348.534250	105.2	–
C ₂ H	4(7/2,4)-3(5/2,3)	349.399342	41.9	2.88
...	4(9/2,4)-3(7/2,3)	349.339067	41.9	2.5
H α	H26 α	353.622747	Rec.	–
HCN	4-3	354.505469	42.5	1.80
HCO+	4-3	356.734250	42.8	1.62