New Periodic Variable Methanol Masers in Massive Star-Forming Regions

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Why are Masers Important?

- Their presence suggest that specific conditions such as density of the environment, temperature in the region, etc.
 - Since they are sensitive species.
- Masers distribution and spectra profiles could suggest possible structure in the region in which they reside.
- Monitoring masers is an indirect way to probe possible changes in the masing region and/or background source.
- The 6.7 and 12.2 GHz methanol masers have one of the rare view of the high mass star forming region



(MASER) is a natural occurring phenomena.



Bartkiewicz et al., 2005



Monitoring Programme for the Methanol masers

- Goedhart et al. (2003,2004) reported periodic variability in seven of 54 monitored at 6.7 GHz.
- Since then, seven more sources were reported by Araya et al. (2010), Szymczak et al. (2011,2015) and Fujiswa et al. (2014) to show periodic variations.



Some of Periodic Variable methanol maser sources



What is the origin of the observed periodicity?

- Sobolev et al. (2007) and Parfenov & Sobolev (2014) proposed variations of dust temperature due to the filaments formed around the young massive stellar object.
- van der Walt, Goedhart & Gaylard (2009) and van der Walt (2011) proposed Colliding Winding Binary (CWB) as the origin of periodicity (Further detail check S.P. van den Heer Poster).
- Araya et al. (2010) proposed periodic circumbinary disc accretion
- Inayoshi et al. (2013) proposed protostellar pulsation.
- There is no observation which confirms any of the above proposal.

The search for the new periodic methanol masers

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Eight sources from the 6.7 GHz Methanol Multibeam Survey Catalogues I and II were selected for monitoring at both 6.7 and 12.2 GHz.

New Periodic methanol masers associated WITH G358.460-0.391



Maswanganye et al. 2015

G358.460-0.391 Maser SPOTS MORPHOLOGY





- Five snapshot observations were made with six 22m Australia Telescope Compact Array (ATCA) antennas.
- The maser morphology is elliptical.
- The major diameter of the fitted ellipse is about 0.26 arcsec.
- With 220 day period and, assuming a total stellar mass of 20 solar mass, the semi-major axes for the two systems will be 1.94 au.

compare waveforms



Group one: the masers show a fairly rapid rise followed by an exponential-like decay to a minimum value {(G9.621+0.196E, G328.24-0.55, G331.13-0.24, from Goedhart et al. (2003, 2004, 2013), G22.357+0.066 (Szymczak et al. 2011), G37.55+0.20 (Araya et al. 2010) and IRAS 22198+6336 (Fujiswa et al. 2014)}



- Group two: The masers show semi sinusoidal type variations {G12.89+0.49 and G188. 95+0.89, from Goedhart et al. (2003, 2004, 2013)}



- Group three: triangular shaped periodic variations (G339.62–0.12, from Goedhart et al. (2003, 2004, 2013))



- Group four: absolute cosine periodic variations (G338.93–0.06, from Goedhart et al. (2003, 2004,2007, 2013) and G358.460-0.391, from Maswanganye et al. (2015)).



G338.93-0.06 AND G358-460-0.391 MINIMUM ANALYSIS



- G338.93–0.06 minimum with good sampling interval.
 - The symmetry could around the minimum could be a suggestions that the origin is a symmetric mechanisms.
- The minimum appears to be not as sharp as expected from the absolute cosine.

- The minimum of G358.460-0.391 looks sharp though it had smaller sampling interval..

MJD

Maswanganye et al.(2015)

Variations in G338.93-0.06 and G358.460-0.391 COULD BE DUE TO AN ECLIPSING BINARY



Hopping Bunny



- Eclipsing binary line curves are similar to that of an absolute cosine (Hopping Bunny).



Could there be more periodic methanol masers?

- Maswanganye et al., (2015) estimate of the probability of a finding a periodic maser in a given sample as 0.13 +/- 0.04.
- It translates into possible 34 +/- 10 possible periodic masers in the 6.7 GHz MMB survey catalogues I, II, III, and IV.
- In the 6.7 GHz MMB survey Catalogues, 262 sources met the selection criteria, 62 had be monitored before, eight were found to be periodic, which suggests 26 or more from the remaining 200 could be periodic.

Ten sources from the 6.7 GHz Methanol Multibeam Survey Catalogues III and IV were selected for monitoring at both 6.7 and 12.2 GHz.



ANOTHER NEW PERIODIC maser



Time delays across the channels



Maswanganye et al., (in Prep)

Trying to explain time lags with maser spots distribution



Maswanganye et al., (in Prep)

Summary

- Origin of the periodicity in class II methanol masers is real and complex to explain (Period range: 29 to 510 days).
- The Waveforms could be a trace of the origin of periodicity.
- There could be 25 or more periodic sources (from 190) in the 6.7 GHz MMB survey catalogues I, II, III, and IV.
- Could there be more waveforms and why are they different ?
- Any mechanism which explain the origin of the periodicity should also explain the diversity of waveforms, unless it is proposed for a particular of methanol masers.