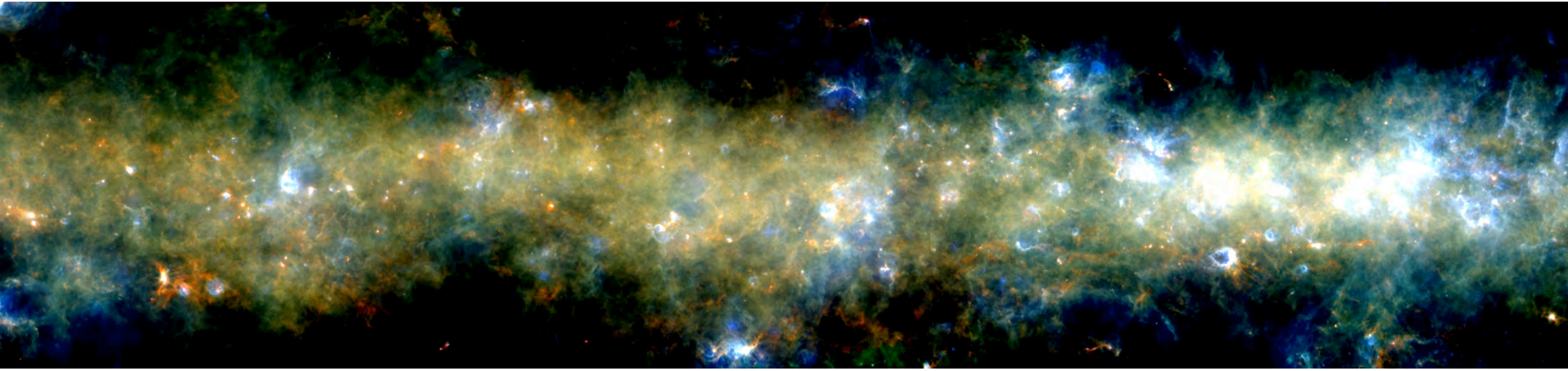


# A Herschel view of massive star formation on the outer Galaxy

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Hi-GAL collaboration

# Thanks to the last generation of Galactic Plane Surveys



## Toward a Predictive Global Model of Galactic Star Formation

- The High-Mass Star Formation Timeline
- Measure the star formation rate and history Galaxy-wide
- Cold dust in the Galactic Plane and the Formation of Molecular Clouds
- Understanding star formation laws and the nature of thresholds as a function of ISM properties across a full range of galactocentric radii metallicity and environmental conditions
- Determining the relative importance of global vs local, spontaneous vs triggering, agents that give rise to star formation.
- Build bottom-up recipes and prescriptions useful for Xgal science

# Hi-GAL

## The Herschel infrared Galactic Plane Survey



Simultaneous 5-bands ( $70\text{-}160\text{-}250\text{-}350\text{-}500\mu\text{m}$ ) continuum mapping of 720 sq. deg. of the Galactic Plane ( $|b| \leq 1^\circ$ ) (Molinari+2010b)

With almost 900 hours observing time is the largest OPEN TIME Herschel KP

**Galaxy-wide Census, Luminosity, Mass and SED of dust structures at all scales from massive YSOs to Spiral Arms**

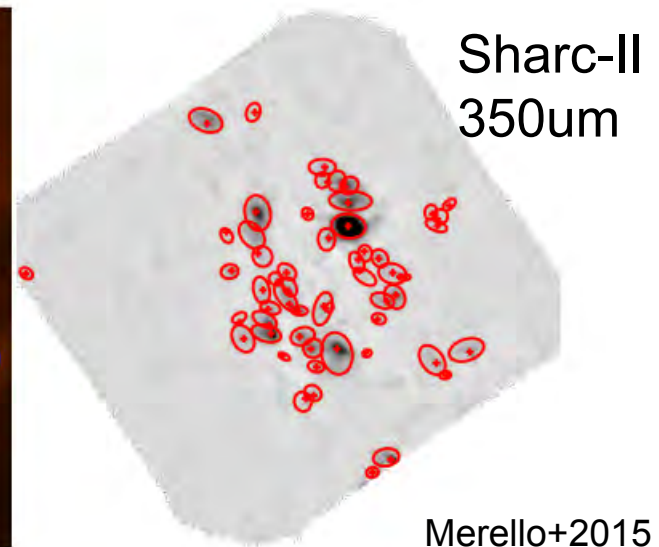
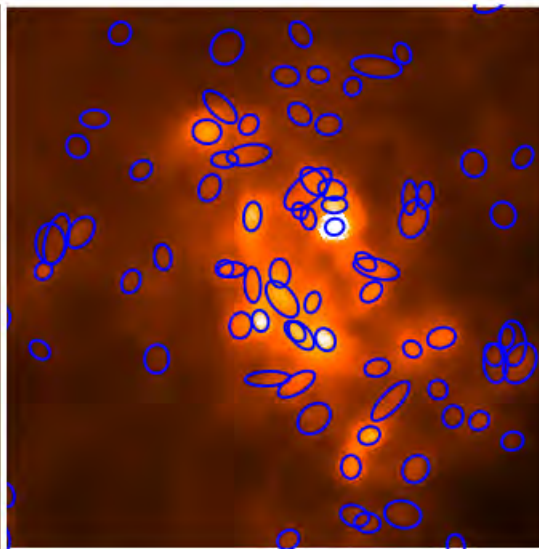
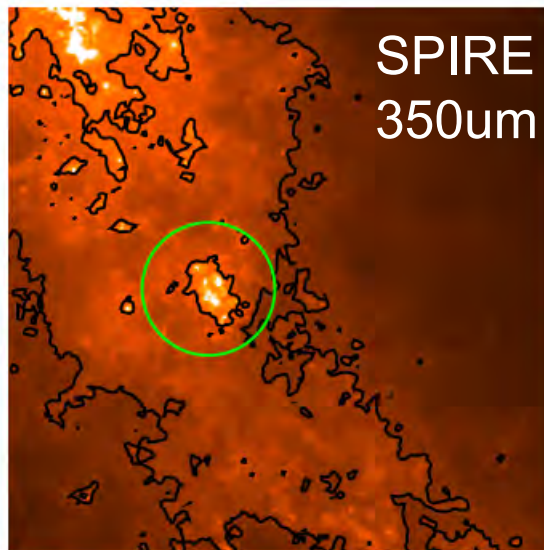
The entire Plane has been observed. Images access (with registered astrometry and absolute flux calibration) and compact source catalogues for longitudes between  $65^\circ$  and  $290^\circ$  (full Inner Galaxy) will come on line soon.



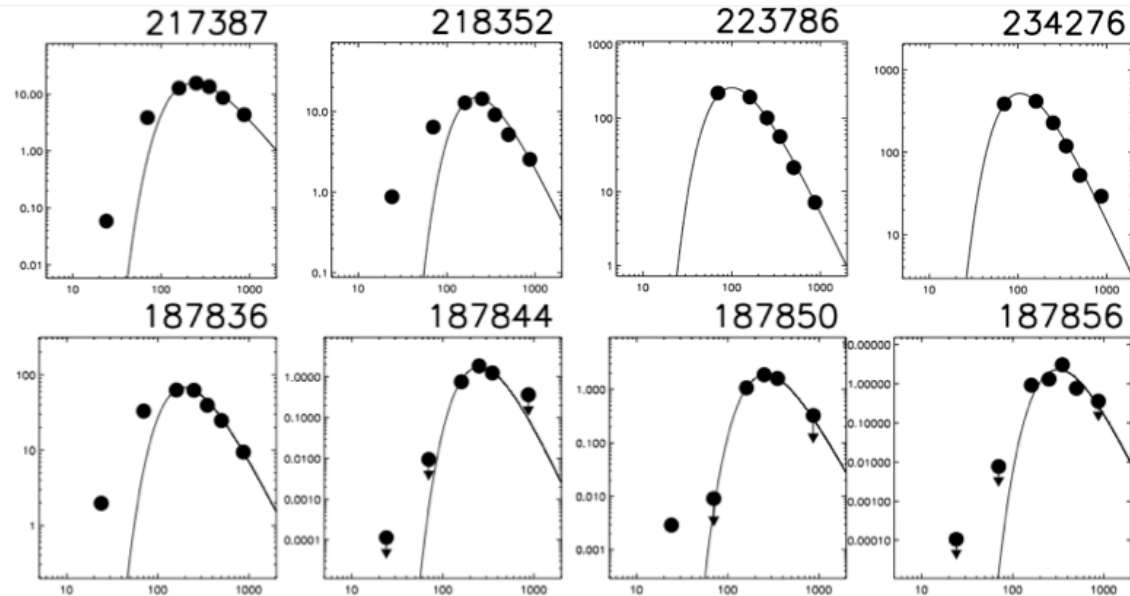
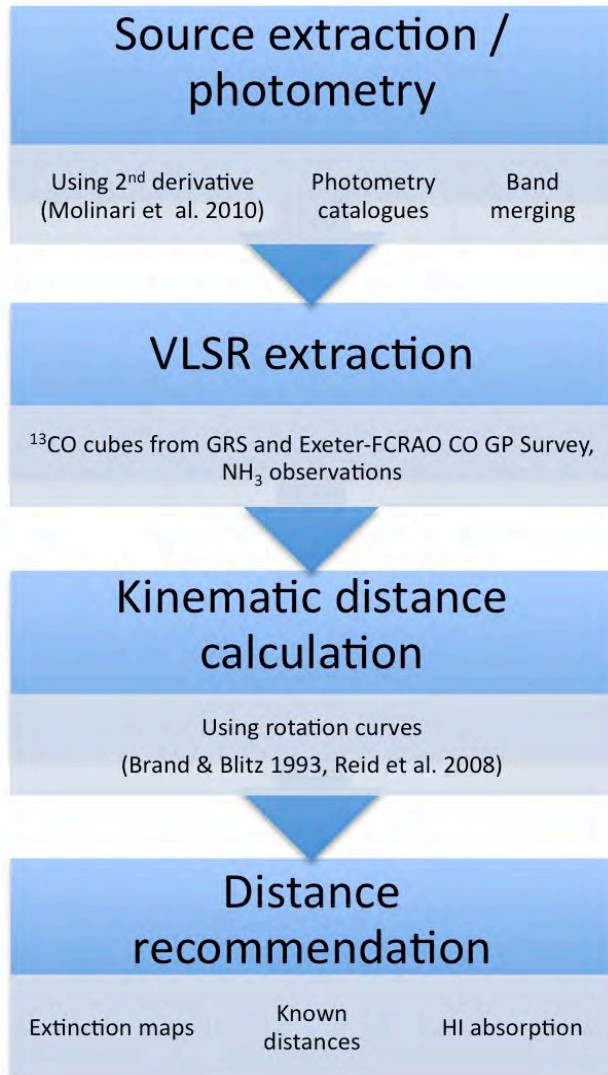
# Hi-GAL is statistics: Huge output...

- First-generation Photometric Catalogues created using CuTEx package (Molinari+11) for the inner Galaxy (done several times....)
- Naïve band-merging produces a catalogue of **519400** entries (Molinari+ 2015, in prep)
- Clump catalogue downselected filtering “nice” SEDs with at least three adjacent counterparts 160-500 $\mu$ m yields **99180** entries

Band	$N_{sources}$
PACS-70 $\mu$ m	122 971
PACS-160 $\mu$ m	292 051
SPIRE-250 $\mu$ m	280 258
SPIRE-350 $\mu$ m	161 855
SPIRE-500 $\mu$ m	85 880



# Hi-GAL is statistics: Huge output...



- For the 99180 inner Galaxy sources with counterparts in at least three bands, we augment SED coverage with ATLASGAL, BGPS, MIPS GAL, WISE, MSX
- A first set of distance estimates (with different levels of reliability) has been carried out excluding  $l < 14^\circ$  and  $l > 350^\circ$ , yielding T, L, M and size, for **56656** sources

Russeil+ 2010

Elia+ 2015, in prep

# Results from Hi-GAL survey: Tips of the Galactic Bar

(Veneziani et al., submitted)

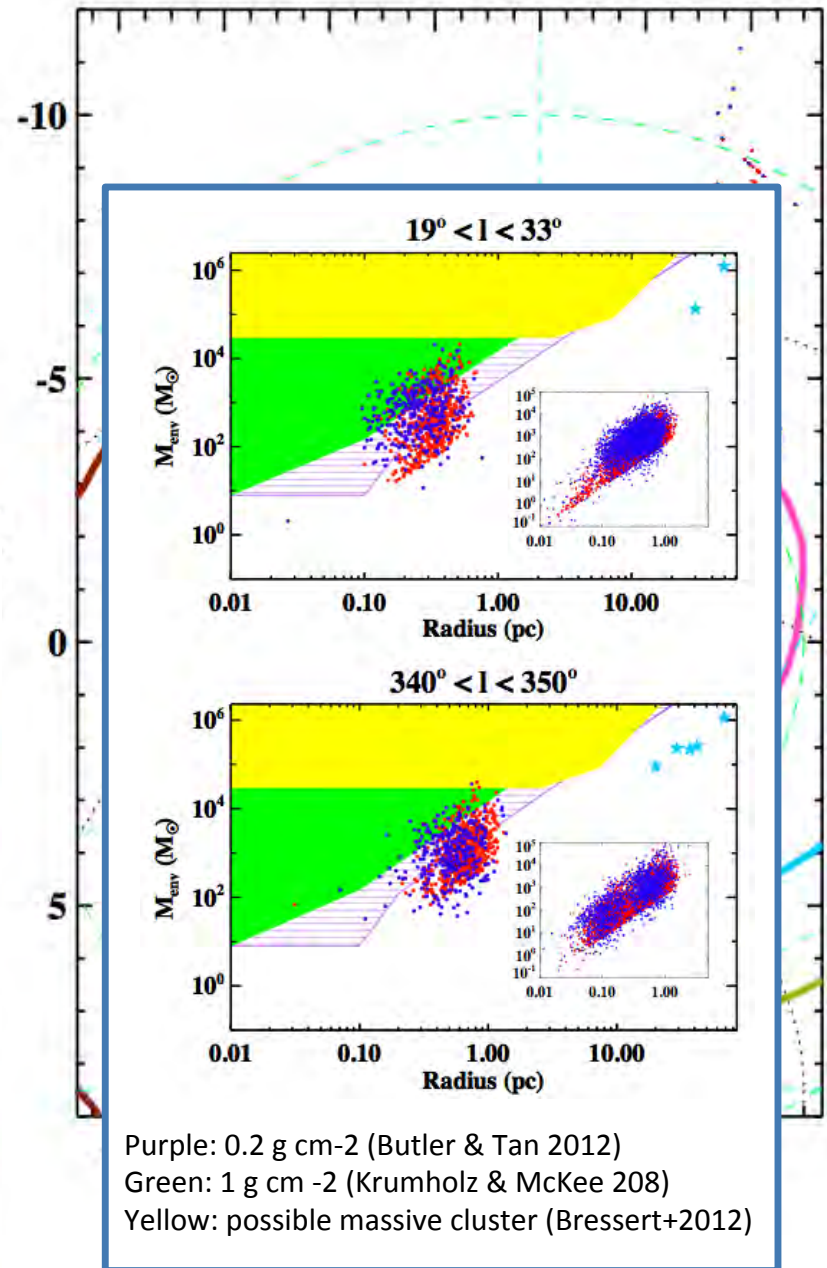
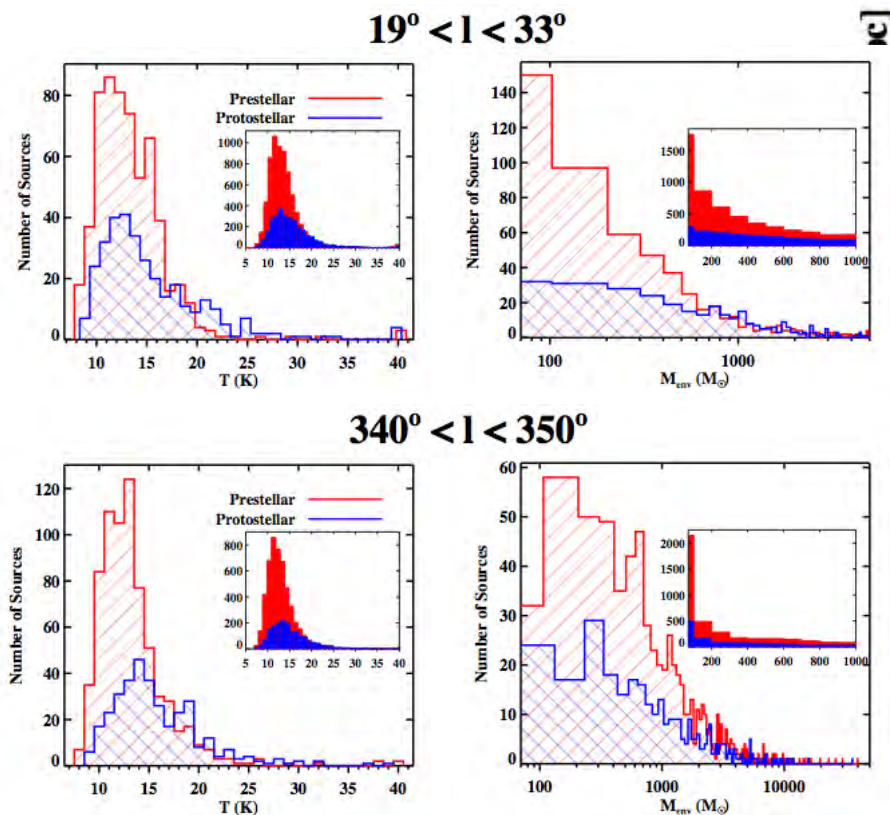
$19 < l < 33$ ,  $340 < l < 350$

Prestellar candidates (without 70  $\mu\text{m}$ )

Protostellar candidates (with 70  $\mu\text{m}$ )

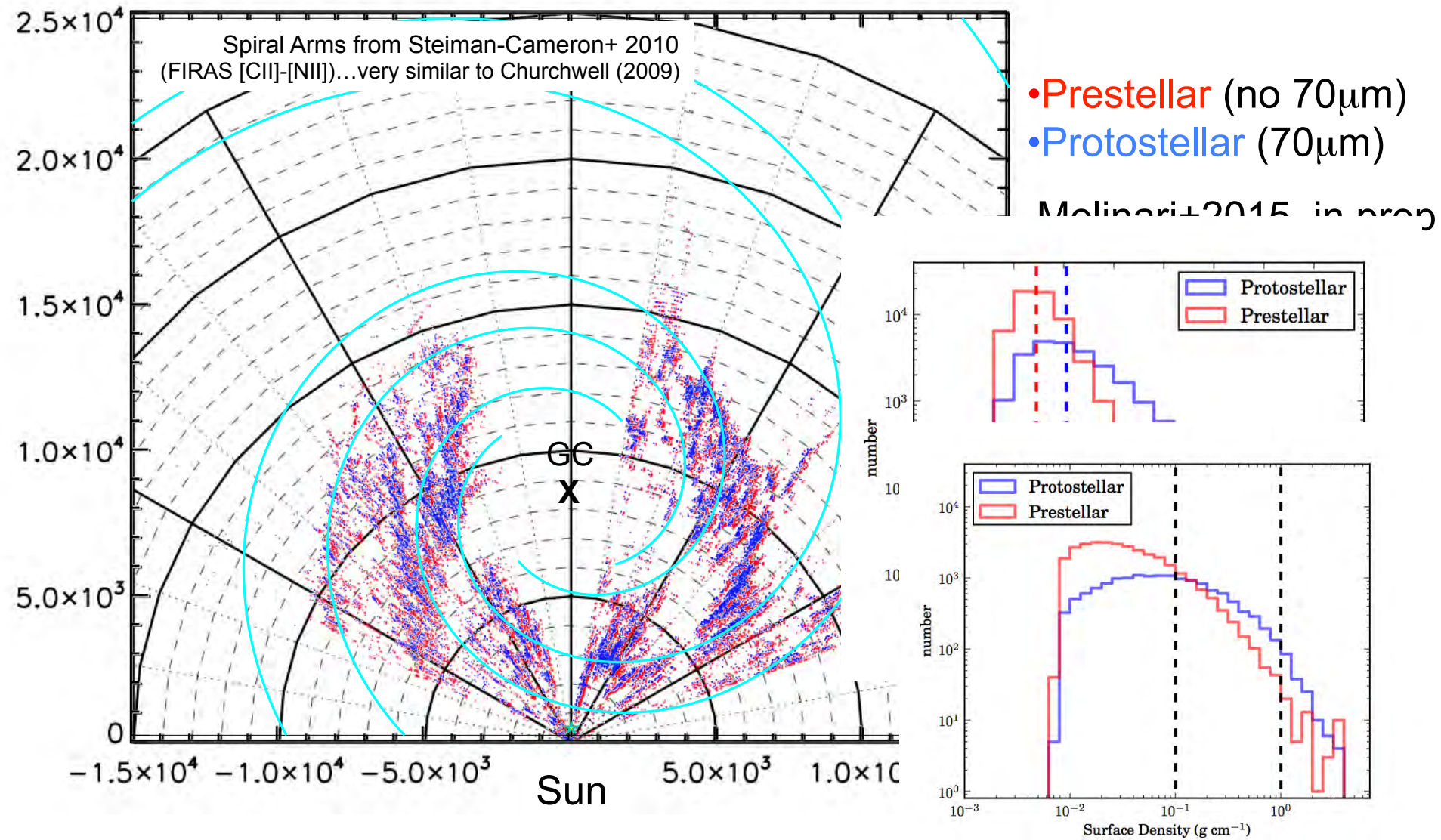
(IV Q) NANTEN 12CO(2-1) (Mizuno & Fukui 2004)

(I Q) FCRAO GRS 13CO(2-1) (Jackson+2006)

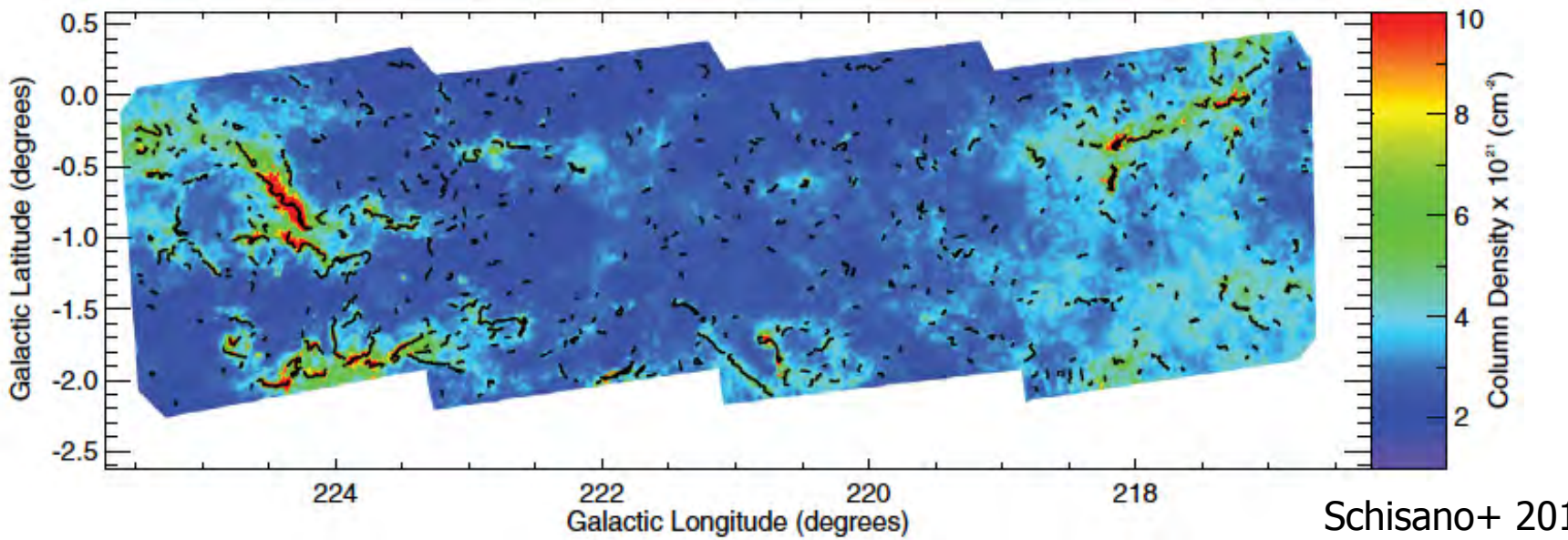
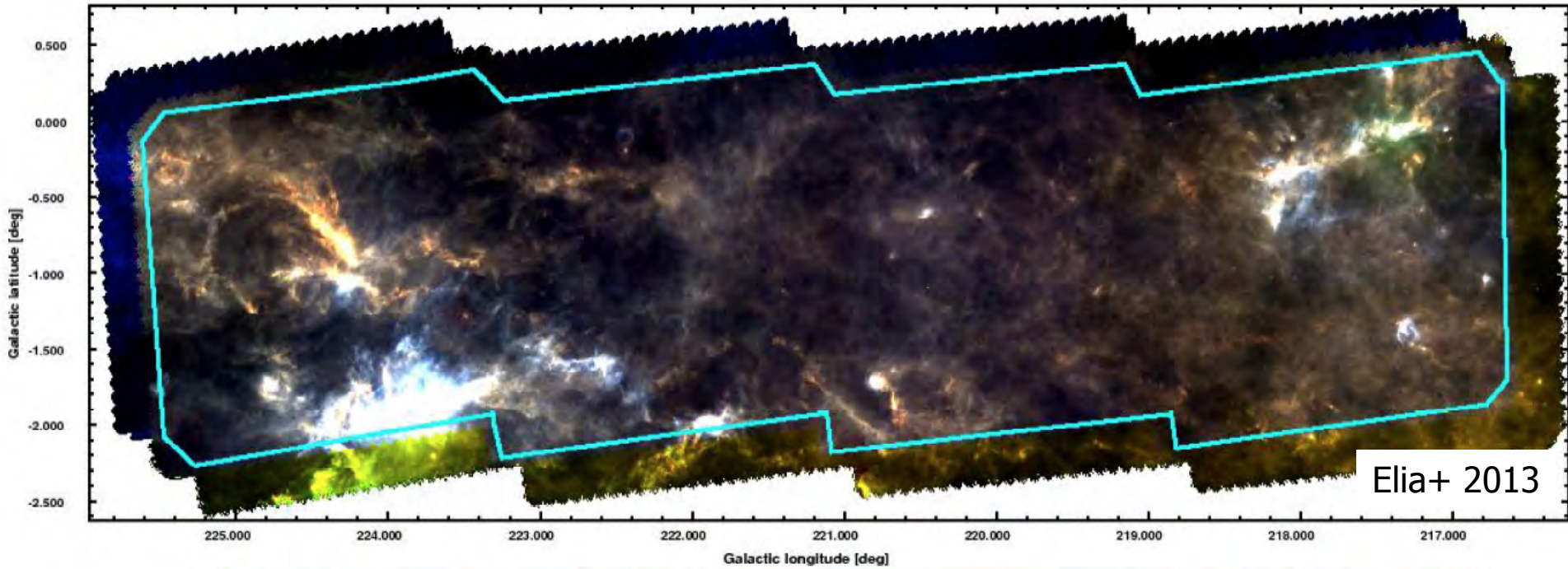




# Global distribution of star formation from Hi-GAL

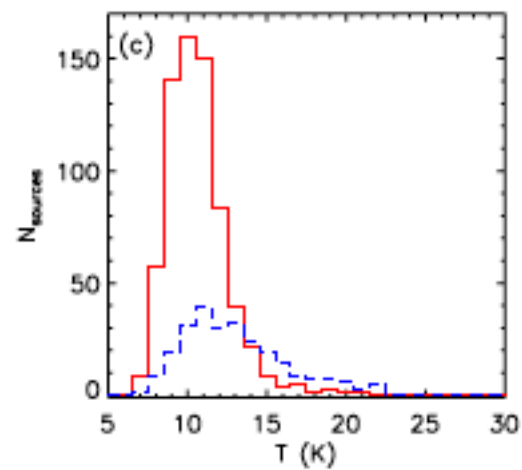
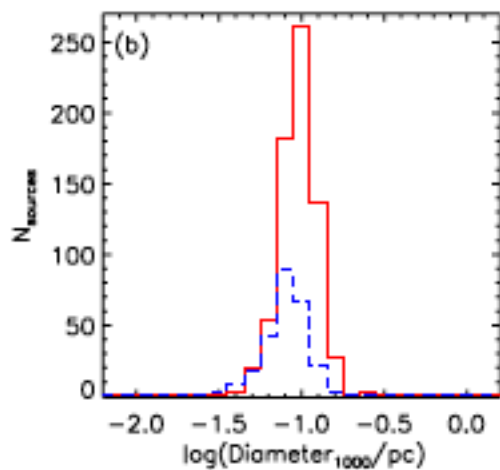
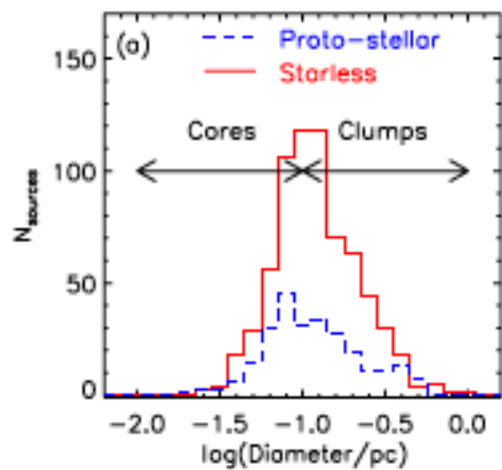
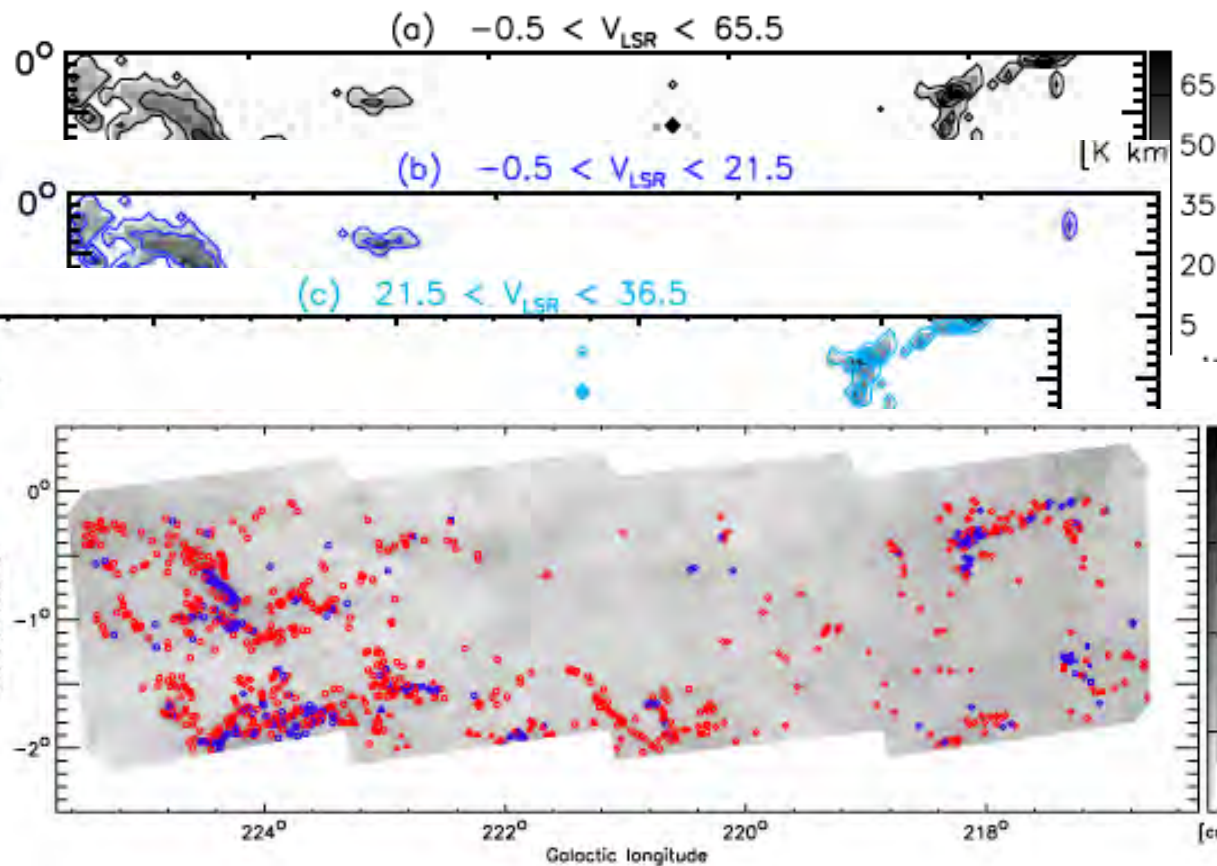
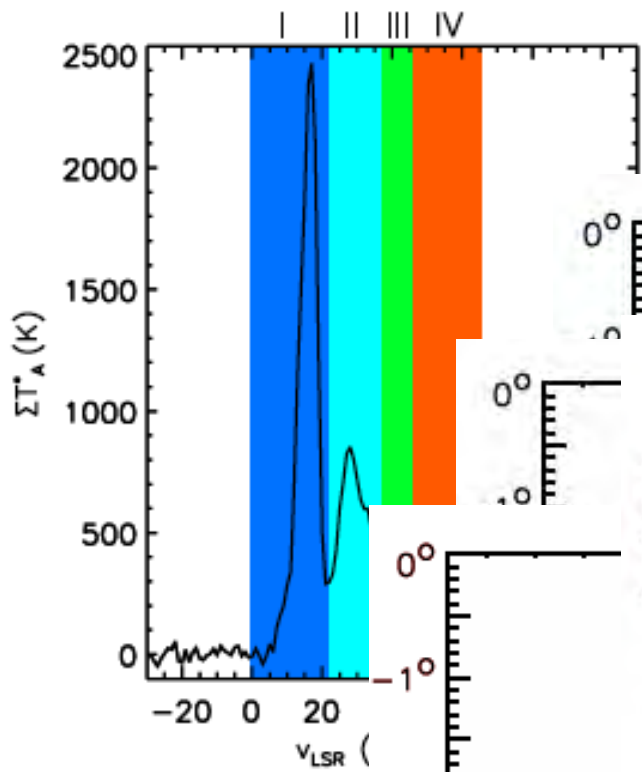


# Herschel view of the outer Galaxy (III Q)

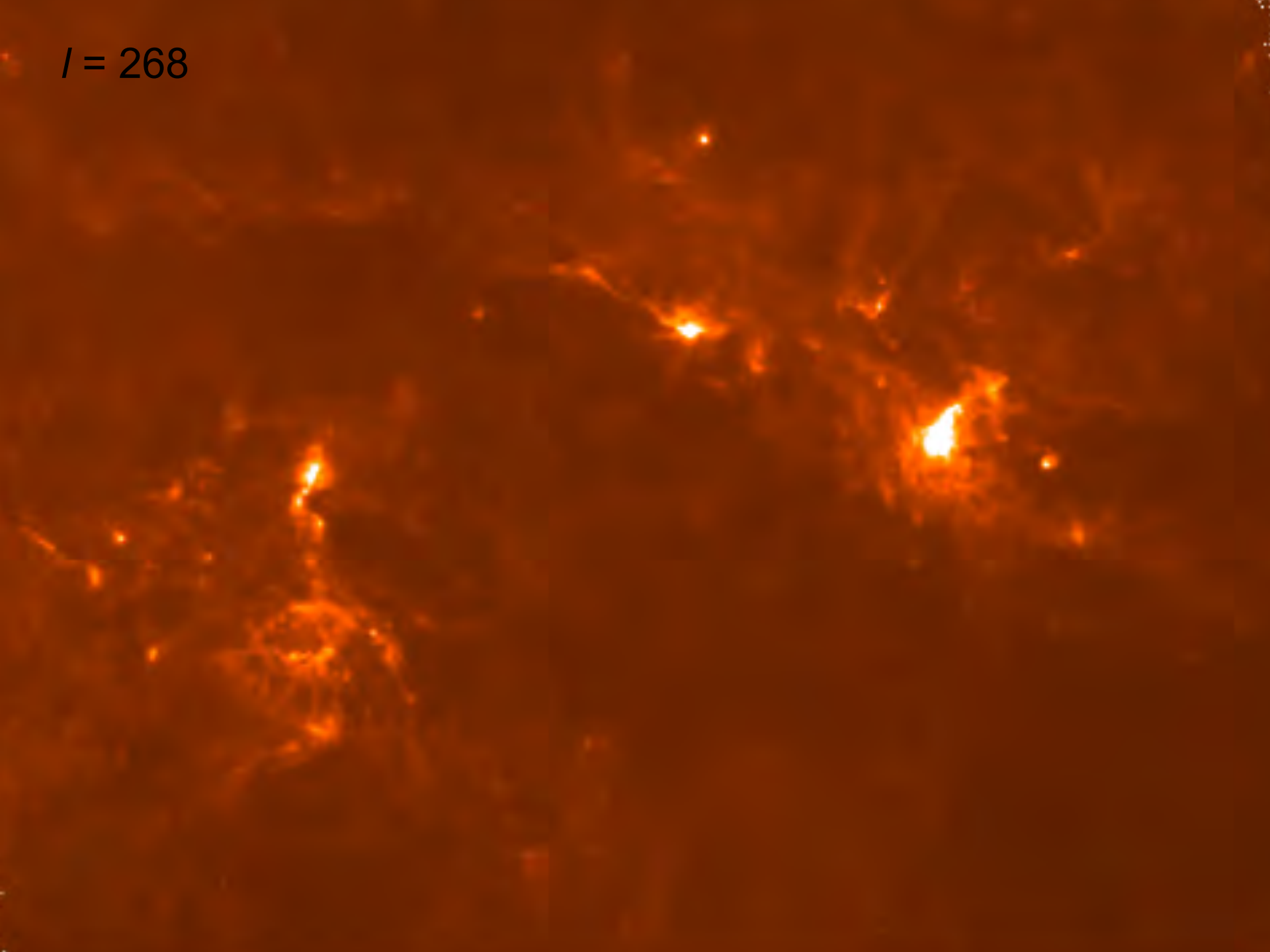




# Distance determination

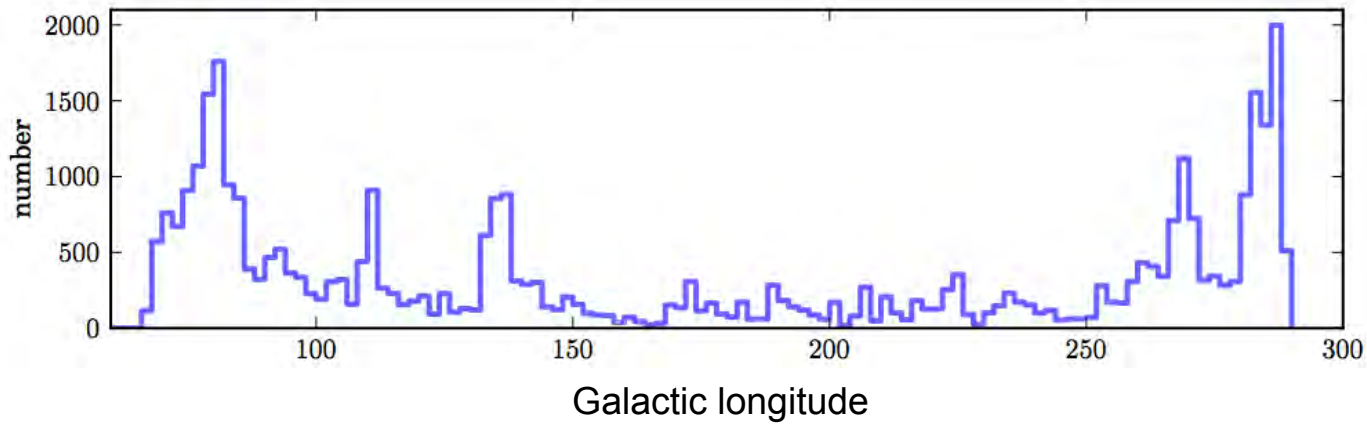


$l = 268$

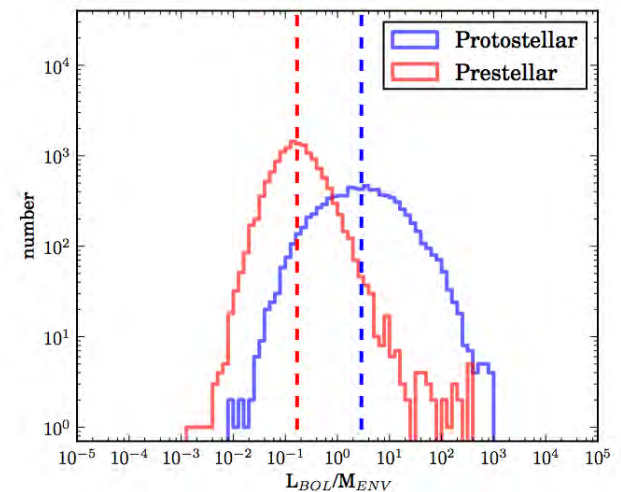
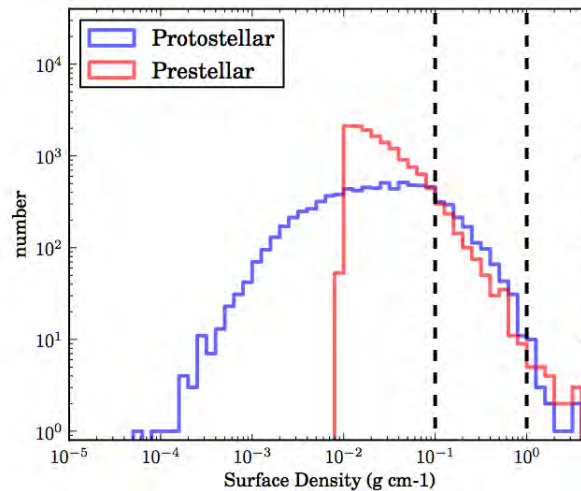
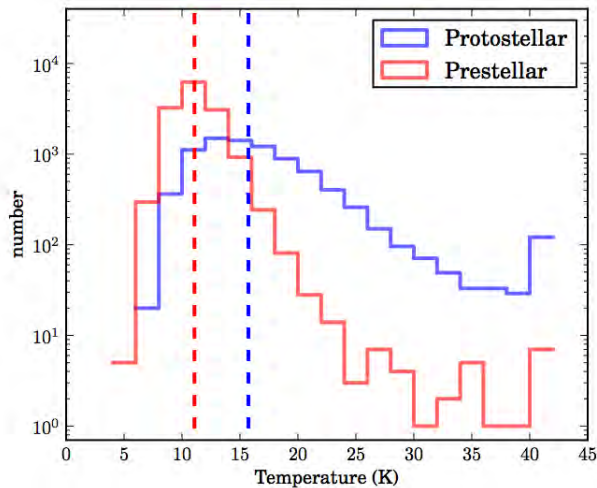




Large number of recovered sources on the outer Galaxy:  
37995 with fitted SED.

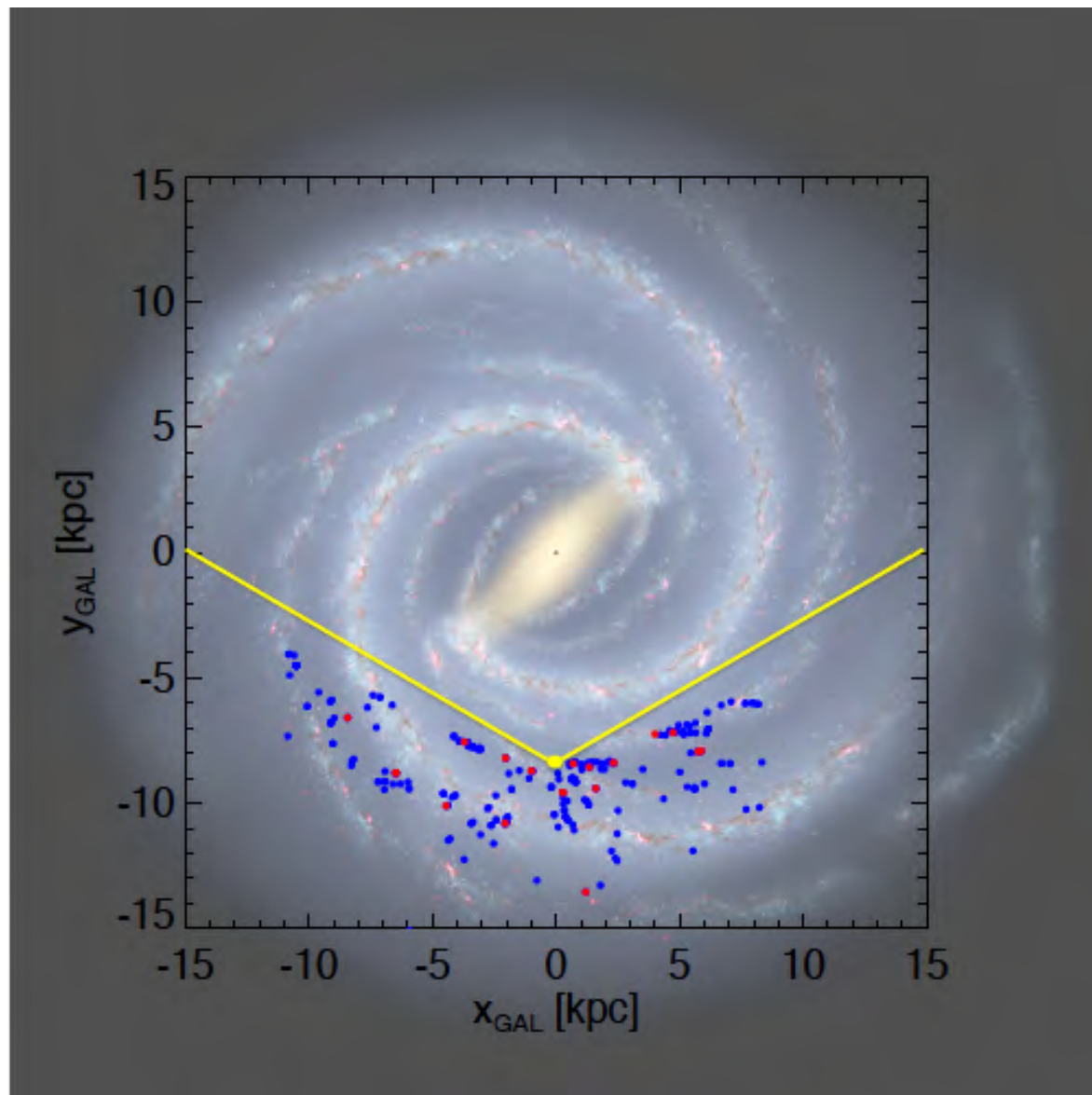


Distance independent parameters:



## Cross-correlation with line surveys:

- Bronfman et al. 1996  
CS(2-1) survey toward  
IRAS points with  
characteristic colors of  
UCHII regions
- New CS(2-1) data  
(Bronfman et al., in prep)
- 291 HiGAL sources with  
CS counterpart.

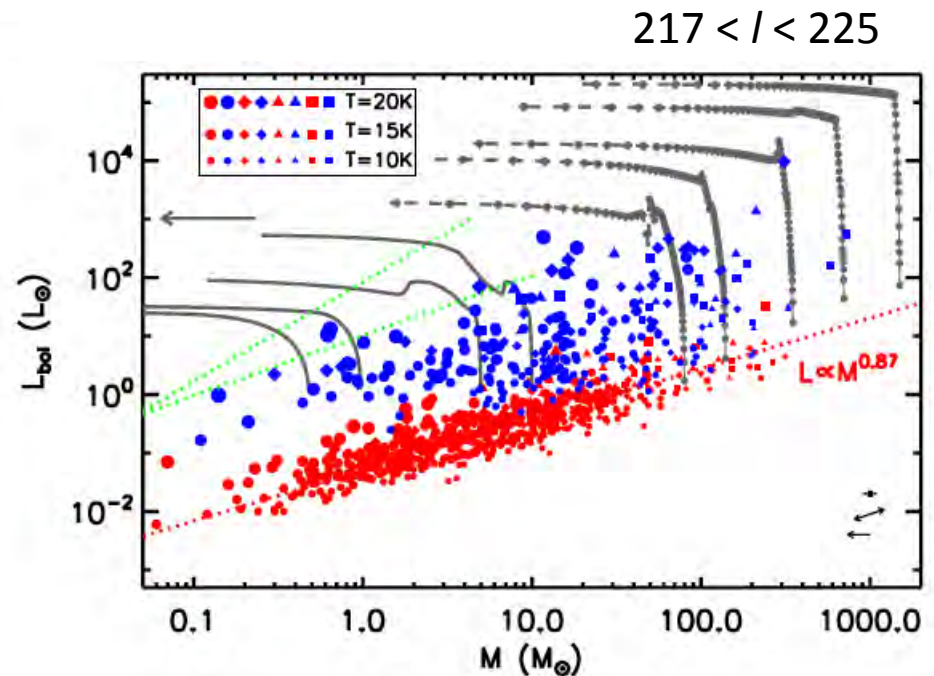
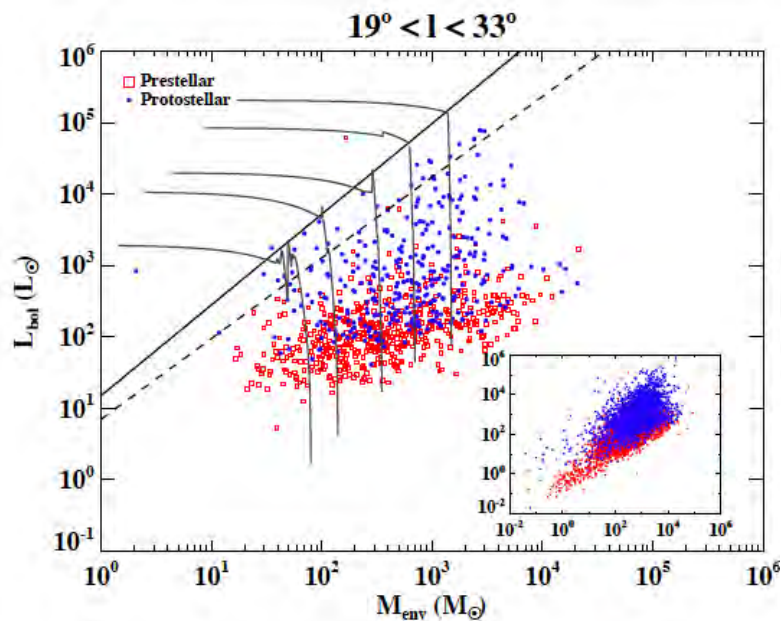






# Lots of things to do:

- Estimation of distances using molecular line catalogs, and distance-dependent properties.
- Characterization of sources in the outer Galaxy, comparison with inner Galaxy.
- Evolutionary scenario for sources
- Filaments in the outer Galaxy and clump association






# Summary

- Hi GAL provides a global picture of Galactic star formation
- 
- Characterization of different evolutionary stages of clumps yields insight on the massive star formation process
- Outer Galaxy maps show population of sources with conditions for massive star formation
- Significant differences between properties of inner and outer Galaxy sources
- Comparison with line surveys allows studies of large scale structure
-

# VIALACTEA



an FP7-SPACE-2013 Collaborative Project



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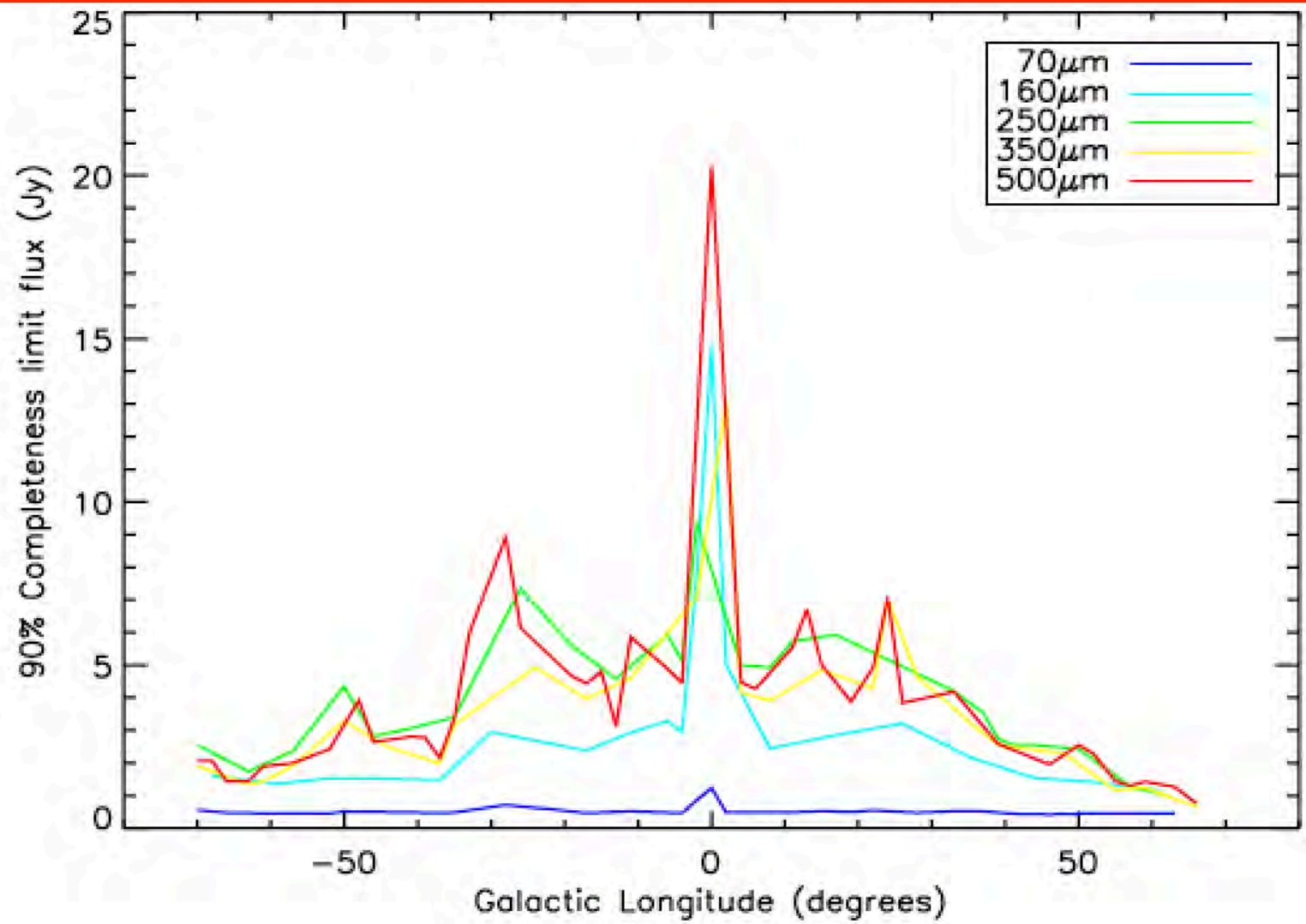


- | **Nagoya Univ.**

The aim is to create a bottom-up model for Galactic star formation, also producing a portal for all community to access all GP Surveys in a unified framework of tools for visualization and science analysis

Thanks





# Comparing L/M

